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## Saddle point approximation to Higher order

Приближение седловой точки к высшему порядку

Апроксимација седласте тачке вишег реда

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## ABSTRACT:

Introduction/purpose: Saddle point approximation has been considered in the paper.

Methods: The saddle point method is used in several different fields of mathematics and physics. Several terms of the expansion for the factorial function have been explicitly computed.

Results: The integrals estimated in this way have values close to the exact one.

Conclusions: Higher order corrections are not negligible even when requiring moderate levels of precision.

KEYWORDS: saddle point approximation, Stirling's formula, Quantum Field Theory..

## Р е з ю м е :

Введение / цель: В данной статье рассмотрено приближение седловой точки.

Методы: Метод седловой точки используется в нескольких различных областях математики и физики. В статье наглядно вычисляются несколько членов расширения для факторной функции.

Результаты: Интегралы, вычисленные таким образом, имеют значения близкие к точному.

Выводы: Поправками высшего порядка не следует пренебрегать, даже в тех случаях, когда требуются умеренные уровни точности.

К л ю ч е в ы е с л о в а : приближение перевала, формула Стирлинга, квантовая теория поля.

## ABSTRACT:

Увод/циљ: У овом раду разматра се апроксимација седласте тачке.

Методе: Метода седласте тачке користи се у неколико различитих области математике и физике. Израчунава се експлицитно неколико чланова проширења за факторску функцију

Резултати: Овако процењени интеграл имају приближно тачне вредности.

Закључак: Корекције вишег реда нису занемариве чак ни када се захтева умерени ниво прецизности.

KEYWORDS: апроксимација седласте тачке, Стирлингова формула, квантна теорија поља.

## AUTHOR NOTES

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## SADDLE POINT METHOD

The saddle point method is an extension of the original method of Laplace (Laplace, 1986) for approximating the value of an integral of the form:

$$\int_a^b \exp[\lambda f(x)] dx, \quad (1)$$

where  $f(x)$  is at least twice differentiable,  $\lambda$  is a large number and the extrema of the integral could also be infinite. Assuming that  $x_0$  is the global maximum of the function  $f(x)$ , Laplace observed that the ratio

$$\frac{\exp[\lambda f(x_0)]}{\exp[\lambda f(x)]} = \exp[\lambda(f(x_0) - f(x))] \quad (2)$$

would increase exponentially with  $\lambda$ , while the ratio

$$\frac{\lambda f(x_0)}{\lambda f(x)} = \frac{f(x_0)}{f(x)} \quad (3)$$

is independent of  $\lambda$ . Therefore, he concluded that the main contribution to the integral (1) comes only from the values of  $x$  in the neighborhood of  $x_0$ , and the latter could be easily calculated.

Our aim is to compute the integral

$$I(\hbar) = \int_{-\infty}^{+\infty} \exp \left[ \frac{if(z)}{\hbar} \right] dz. \quad (4)$$

Following the notation of (Parisi, 1988), we expand the so-called saddle point approximation first proposed by Daniels (Daniels, 1954) (also known as the steepest descend method) beyond first order approximation obtaining several terms of approximation, which is the main scope of this paper. As usual, one expands about the maximum  $df/dz = 0$  obtaining a Gaussian integral for  $I(\#)$ , e.g. as in the Stirling's formula for  $n!$ . This suffices for many applications, as the Gaussian falls down quite quickly so further corrections are usually not necessary, unless a precision better than the percent order is required as it will be seen.

We want to compute eq. (4) beyond the first order in  $\#$ . From here onward,  $\#$  plays a role of a generic small expansion parameter beyond its physical meaning. In order to achieve this goal, we expand  $f(z)$  around the critical point  $z_0$  such that  $df(z_0)/dz = 0$ :

$$f(z) = f(z_0) + \frac{1}{2}f^{(2)}(z_0)(z-z_0)^2 + \frac{1}{6}f^{(3)}(z_0)(z-z_0)^3 + \frac{1}{24}f^{(4)}(z_0)(z-z_0)^4 + \mathcal{O}((z-z_0)^5)$$

(5)

The trick is to separate the exponential in two parts: the Gaussian and the remnant. The latter is expanded again in Taylor's series, i.e. we write:

$$\begin{aligned} \exp \left[ \frac{if(z)}{\hbar} \right] &= \exp \left[ \frac{i(f(z_0) + \frac{1}{2}f^{(2)}(z_0)(z - z_0)^2)}{\hbar} \right] \\ &\times \exp \left[ \frac{i}{\hbar} \left( \frac{1}{6}f^{(3)}(z_0)(z - z_0)^3 + \frac{1}{24}f^{(4)}(z_0)(z - z_0)^4 + \mathcal{O}((z - z_0)^5) \right) \right] \end{aligned} \quad (6)$$

that is, a Gaussian times some other function that will be eventually expanded in Taylor's series. We could rewrite eq. (6) as

$$\exp \left[ \frac{if(z)}{\hbar} \right] = \exp \left[ \frac{i(f(z_0) + \frac{1}{2}f^{(2)}(z_0)(z - z_0)^2)}{\hbar} \right] \times \exp \left[ \frac{ig(z)}{\hbar} \right] \quad (7)$$

where at least formally  $g(z)$  is the remainder from the third order of the expansion of  $f(z)$ :

$$g(z) = \sum_{n=3}^{+\infty} \frac{f^{(n)}(z_0)}{n!} (z - z_0)^n \quad (8)$$

Of course Taylor's expansion of eq. (8) is not the one of  $f(z)$  given in eq. (5) due to the exponential function. Great care has to be applied in order to pick the right power of  $\hbar$ . For instance, to second order in  $\hbar$  we have:

$$\begin{aligned} \exp \left[ \frac{ig(z)}{\hbar} \right] &= \exp \left[ \frac{ig(z_0)}{\hbar} \right] \times \left[ 1 + \frac{i}{\hbar} g'(z_0)(z - z_0) \right] \\ &+ \exp \left[ \frac{ig(z_0)}{\hbar} \right] \times \left[ \frac{1}{2\hbar^2} (i\hbar g''(z_0) - g'(z_0)^2) (z - z_0)^2 \right] \end{aligned} \quad (9)$$

and powers of  $\hbar$  are mixed as it can be seen. We obtain

$$\exp \left[ \frac{ig(z)}{\hbar} \right] = \sum_{n=0}^{+\infty} \frac{\phi^{(n)}(z_0)}{n!} (z - z_0)^n \quad (10)$$

for  $\phi(z) = \exp(ig(z)/\hbar)$ . Plugging it back in eqs (7) and (4), we obtain

$$I(\hbar) = \exp \left[ \frac{if(z_0)}{\hbar} \right] \int_{-\infty}^{+\infty} \exp \left[ \frac{if^{(2)}(z_0)(z-z_0)^2}{2\hbar} \right] \times \sum_{n=0}^{+\infty} \frac{\phi^{(n)}(z_0)}{n!} (z-z_0)^n dz \quad (11)$$

Pulling the sum out of the integral shows clearly that only even powers survive because of the Gaussian integral.

Calling  $I_0$  the Gaussian integral

$$I_0(\hbar) = \exp \left[ \frac{if(z_0)}{\hbar} \right] \int_{-\infty}^{+\infty} \exp \left[ \frac{if^{(2)}(z_0)(z-z_0)^2}{2\hbar} \right] dz \quad (12)$$

that has the value

$$I_0(\hbar) = \exp \left[ \frac{if(z_0)}{\hbar} \right] \left[ \frac{2\pi i \hbar}{f^{(2)}(z_0)} \right]^{1/2} \quad (13)$$

compared to eq. (4) gives the result to first order in  $\hbar$

$$I(\hbar) = I_0(\hbar)(1 + \mathcal{O}(\hbar)) \quad (14)$$

With a notation where  $f^{(n)}$  is the  $n$ th derivative of  $f(z)$  computed in  $z_0$ , the  $\mathcal{O}(\hbar^2)$  correction to  $I(\hbar)$  is given by:

$$I_2(\hbar) = \frac{5(f^{(3)})^2 - 3f^{(2)}f^{(4)}}{24(f^{(2)})^3} \quad (15)$$

while the  $\mathcal{O}(\hbar^3)$  correction reads

$$I_3(\hbar) = \frac{-24(f^{(2)})^3 f^{(6)} + (f^{(2)})^2 (168 f^{(3)} f^{(5)} + 105 (f^{(4)})^2)}{1152 (f^{(2)})^6} - \frac{630 f^{(2)} (f^{(3)})^2 f^{(4)} + 385 (f^{(3)})^4}{1152 (f^{(2)})^6}. \quad (16)$$

That is

$$I(\hbar) = I_0(\hbar) [1 + (i\hbar)I_2(\hbar) + (i\hbar)^2 I_3(\hbar) + \mathcal{O}(\hbar^3)] . \quad (17)$$

More terms of the expansion have been calculated and terms up to  $\mathcal{O}(\hbar^7)$  are shown in the Appendix.

This kind of approximation is often used in physics, in statistical mechanics when counting the configurations by means of Stirling's formula (see later). The WKB approximation can be thought of as a saddle point approximation (Wentzel, 1926; Kramers, 1926; Brillouin, 1926). Starting from the work of Dirac (Dirac, 1933), Feynman devised the method of the path integral and with a saddle point approximation derived the Schrödinger equation (Feynman, 1965)

In the quantum field theory, for example, it is used to evaluate path integral perturbatively in order to compute the effective action for a given model (Ramond, 1989). Consider for instance the action  $S$  of a bosonic field  $\phi$ :

$$S[\varphi] = \int \frac{1}{2}(\partial\varphi)^2 + \frac{m}{2}\varphi^2 + V(\varphi) d^4x. \quad (18)$$

One could then apply the procedure of eq. (11), expanding the path integral in the Euclidean space around the classical field  $\phi_0$  which is extremal for the action (18), i.e.

$$\left. \frac{\delta S[\varphi]}{\delta \varphi} \right|_{\varphi=\varphi_0} = 0 \quad (19)$$

and performing the Gaussian integral yields the standard result:

$$\Gamma[\varphi] = S[\varphi_0] + \frac{\hbar}{2} \text{Tr} [\log(-\partial^2 + m^2 + V''(\varphi_0))] + \mathcal{O}(\hbar^2). \quad (20)$$

Including more terms in the expression beyond the leading order of eq. (13) shows that the resulting analytic approximation retains its validity over the whole integration range, not just towards the point  $z_0$ .

#### AN EXAMPLE: STIRLING'S APPROXIMATION

The expression given in eq. (17) has been verified with Stirling's formula (Stirling, 1764) for the Gamma function, given by

$$\Gamma(z+1) = \int_0^{+\infty} t^z \exp(-t) dt = \int_0^{+\infty} \exp(-t + z \log(t)) dt \quad (21)$$

which is equal to  $n!$  when  $z$  is an integer  $n$ . With the position  $\# = -i$  and  $f(t) = t - z \log(t)$  using the formulae starting from expansion of eq. (17) and considering the terms given in eqs. (23)–(26), we obtain the fifth order for  $z \rightarrow +\infty$ :

$$\Gamma(z+1) = \sqrt{2\pi z} \left(\frac{z}{e}\right)^z \left[ 1 + \frac{1}{12z} + \frac{1}{288z^2} - \frac{139}{51840z^3} - \frac{571}{2488320z^4} + \mathcal{O}\left(\frac{1}{z^5}\right) \right]. \quad (22)$$

After the publication of the book of de Moivre (Moivre, 1730) where he developed an approximation to  $(\frac{n}{n/2})^{n/2}$  while developing general procedures for probability, Stirling found his asymptotic series

(22) for  $\log n!$  improving de Moivre's result and introducing the "Stirling's constant"  $(\log 2\pi)/2$ . After this result, de Moivre used a different method to compute the asymptotic series to  $\log n!$  obtaining a similar expansion (Moivre, 1730, 1756).

Notice that Stirling's asymptotic expansion of eq. (22) is not a convergent series (Whittaker & Watson, 1927; Erdelyi, 1956), that is, at the fixed  $z$  the accuracy improves when adding more terms, up to a point where it actually gets worse while increasing the approximation order.

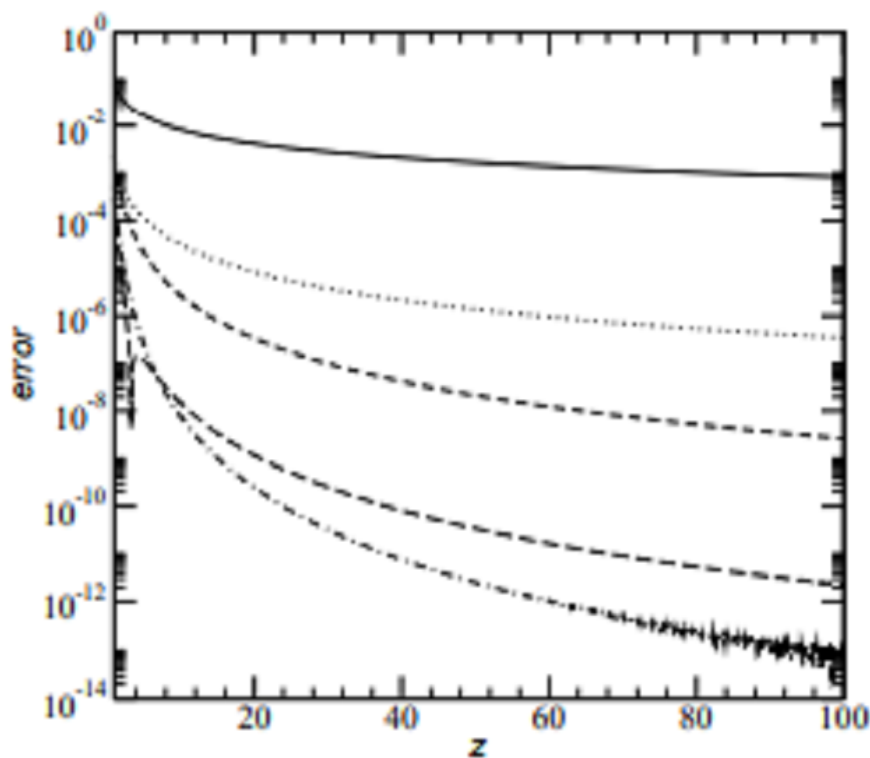


FIGURE 1

Relative error for Stirling's approximation of  $\Gamma(z)$  as a function of  $z$ . The various decreasing curves are in the increasing approximation order, from 1 to 5 terms.

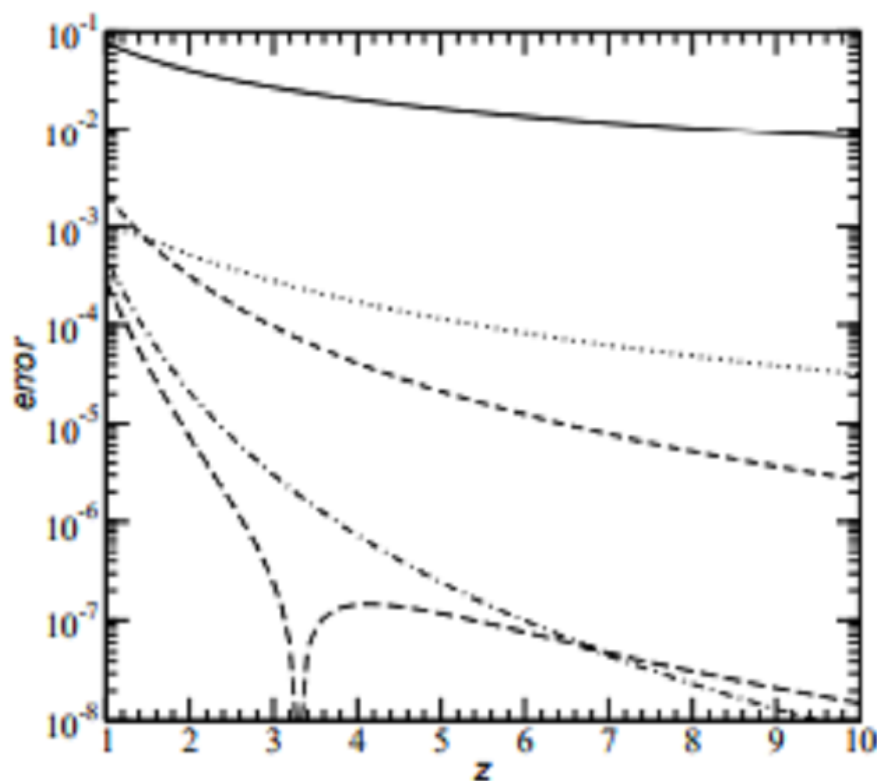


FIGURE 2

The same plot as Fig. (1) for a positive  $z$  range less than 10. This enhancement shows the crossing of the accuracy for various approximation orders.

In Fig. (1), we have shown the relative error of the first 5 terms approximating  $\Gamma(z)$  as the functions of  $z$ . As it could be seen, the increasing order shows better accuracy for the values of  $z$  larger than about 10, as one could expect from the structure of eq. (22).

One could readily notice that the first order approximation is not enough if requiring a better accuracy than one at the percentage level. From Fig. (2), it is also clear than for a small  $z$  a great level of accuracy could only be met by retaining several orders of approximation.

In Table 1, we show some values of  $n!$  for small values of  $n$  and compare the results of different approximation orders. It readily appears that, even to achieve the precision of a pocket calculator, we have to retain several terms of eq. (17), and in particular for those  $n$  one has to consider at least the one shown in eq. (25), much more complicated than the simple expression usually cited of eq. (13).

TABLE 1

The value of  $n!$  for different orders of approximation

$n$	$n!$	$\mathcal{O}(\hbar)$	$\mathcal{O}(\hbar^2)$	$\mathcal{O}(\hbar^3)$	$\mathcal{O}(\hbar^4)$
9	362880	359536.9	362865.9	362881.3	362880.0
10	3628800	3598695.6	3628684.7	3628809.7	3628800.0
11	39916800	39615625.1	39915743.4	39916880.2	39916800.4
12	479001600	475687486.5	478990871.8	479002341.9	479001603.7



## CONCLUSIONS

We have shown in some detail the procedure of computing the integrals via the saddle point method, also known as the steepest descent method, which finds its application in several branches ranging from theoretical physics to computational methods. We have explicitly computed many terms of this asymptotic expansions furnishing analytical results, and applied its results to a wellknown integral, estimating the error. We have also shown that, in order to obtain a certain degree of precision, the usual Gaussian term is not enough and a better approximation should be pursued.

## APPENDIX

Here  $f_n$  refers to the  $n$ th derivative of  $f$  taken at the point  $z_0$ .

**Second order  $\mathcal{O}(\#^2)$ :**

$$\frac{5 f_3^2}{24 f_2^3} - \frac{f_4}{8 f_2^2} \quad (23)$$

**Third order  $\mathcal{O}(\#^3)$ :**

$$-\frac{f_6}{48 f_2^3} + \frac{56 f_3 f_5 + 35 f_4^2}{384 f_2^4} - \frac{35 f_3^2 f_4}{64 f_2^5} + \frac{385 f_3^4}{1152 f_2^6} \quad (24)$$

**Fourth order  $\mathcal{O}(\#^4)$ :**

$$\begin{aligned} & -\frac{f_8}{384 f_2^4} - \frac{-20 f_3 f_7 - 35 f_4 f_6 - 21 f_5^2}{640 f_2^5} + \frac{-616 f_3^2 f_6 - 1848 f_3 f_4 f_5}{3072 f_2^6} \\ & - \frac{385 f_4^3}{3072 f_2^6} - \frac{-8008 f_3^3 f_5 - 15015 f_3^2 f_4^2}{9216 f_2^7} - \frac{25025 f_3^4 f_4}{9216 f_2^8} + \frac{85085 f_3^6}{82944 f_2^9} \end{aligned} \quad (25)$$

**Fifth order  $\mathcal{O}(\#^5)$ :**

$$\begin{aligned}
& -\frac{f_{10}}{3840 f_2^5} + \frac{220 f_3 f_9 + 495 f_4 f_8 + 792 f_5 f_7 + 462 f_6^2}{46080 f_2^6} \\
& + \frac{-1430 f_3^2 f_8 - 5720 f_3 f_4 f_7 - (8008 f_3 f_5 + 5005 f_4^2) f_6 - 6006 f_4 f_5^2}{30720 f_2^7} \\
& + \frac{91520 f_3^3 f_7 + 480480 f_3^2 f_4 f_6 + 288288 f_3^2 f_5^2 + 720720 f_3 f_4^2 f_5 + 75075 f_4^4}{294912 f_2^8} \\
& + \frac{-340340 f_3^4 f_6 - 2042040 f_3^3 f_4 f_5 - 1276275 f_3^2 f_4^3}{221184 f_2^9} \\
& + \frac{2586584 f_3^5 f_5 + 8083075 f_3^4 f_4^2}{442368 f_2^{10}} - \frac{11316305 f_3^6 f_4}{663552 f_2^{11}} + \frac{37182145 f_3^8}{7962624 f_2^{12}}
\end{aligned} \tag{26}$$

Sixth order  $\mathcal{O}(\#^6)$ :

$$\begin{aligned}
& -\frac{f_{12}}{46080 f_2^6} - \frac{-364 f_3 f_{11} - 1001 f_4 f_{10} - 2002 f_5 f_9 - 3003 f_6 f_8 - 1716 f_7^2}{645120 f_2^7} \\
& + \frac{-5720 f_3^2 f_{10} - 28600 f_3 f_4 f_9 - (51480 f_3 f_5 + 32175 f_4^2) f_8}{737280 f_2^8}
\end{aligned} \tag{27a}$$

$$\begin{aligned}
& - \frac{(68640 f_3 f_6 + 102960 f_4 f_5) f_7 - 60060 f_4 f_6^2 - 72072 f_5^2 f_6}{737280 f_2^8} \\
& - \frac{-486200 f_3^3 f_9 - 3281850 f_3^2 f_4 f_8 - (5250960 f_3^2 f_5 + 6563700 f_3 f_4^2) f_7}{6635520 f_2^9} \\
& - \frac{3063060 f_3^2 f_6^2 - (18378360 f_3 f_4 f_5 + 3828825 f_4^3) f_6 - 3675672 f_3 f_5^3}{6635520 f_2^9} \\
& - \frac{6891885 f_4^2 f_5^2}{6635520 f_2^9} \\
& + \frac{-3695120 f_3^4 f_8 - 29560960 f_3^3 f_4 f_7 - (41385344 f_3^3 f_5 + 77597520 f_3^2 f_4^2) f_6}{7077888 f_2^{10}} \\
& - \frac{93117024 f_3^2 f_4 f_5^2 - 77597520 f_3 f_4^3 f_5 - 4849845 f_4^5}{7077888 f_2^{10}} \\
& - \frac{-20692672 f_3^5 f_7 - 181060880 f_3^4 f_4 f_6 - 108636528 f_3^4 f_5^2}{7077888 f_2^{11}} \\
& - \frac{543182640 f_3^3 f_4^2 f_5}{7077888 f_2^{11}} \\
& - \frac{169744575 f_3^2 f_4^4}{7077888 f_2^{11}} + \frac{-416440024 f_3^6 f_6 - 3747960216 f_3^5 f_4 f_5}{31850496 f_2^{12}} \\
& - \frac{3904125225 f_3^4 f_4^3}{31850496 f_2^{12}} \\
& - \frac{-1487285800 f_3^7 f_5 - 6506875375 f_3^6 f_4^2}{31850496 f_2^{13}} - \frac{929553625 f_3^8 f_4}{7077888 f_2^{14}} + \frac{5391411025 f_3^{10}}{191102976 f_2^{15}}
\end{aligned} \tag{27b}$$

Seventh order  $\phi^{(7)}$ :

$$\begin{aligned}
 & -\frac{f_{14}}{645120 f_2^7} + \frac{560 f_3 f_{13} + 1820 f_4 f_{12} + 4368 f_5 f_{11} + 8008 f_6 f_{10}}{10321920 f_2^8} \\
 & \quad + \frac{11440 f_7 f_9 + 6435 f_8^2}{10321920 f_2^8} \\
 & + \frac{-30940 f_3^2 f_{12} - 185640 f_3 f_4 f_{11} - (408408 f_3 f_5 + 255255 f_4^2) f_{10}}{30965760 f_2^9} \\
 & \quad - \frac{(680680 f_3 f_6 + 1021020 f_4 f_5) f_9}{30965760 f_2^9}
 \end{aligned} \tag{28a}$$

$$\begin{aligned}
 & - \frac{(875160 f_3 f_7 + 1531530 f_4 f_6 + 918918 f_5^2) f_8}{30965760 f_2^9} \\
 & - \frac{875160 f_4 f_7^2 - 2450448 f_5 f_6 f_7 - 476476 f_6^3}{30965760 f_2^9} \\
 & + \frac{23514400 f_3^3 f_{11} + 193993800 f_3^2 f_4 f_{10}}{1857945600 f_2^{10}} \\
 & + \frac{(387987600 f_3^2 f_5 + 484984500 f_3 f_4^2) f_9}{1857945600 f_2^{10}} \\
 & + \frac{(581981400 f_3^2 f_6 + 1745944200 f_3 f_4 f_5 + 363738375 f_4^3) f_8}{1857945600 f_2^{10}} \\
 & \quad + \frac{332560800 f_3^2 f_7^2}{1857945600 f_2^{10}} \\
 & + \frac{(2327925600 f_3 f_4 f_6 + 1396755360 f_3 f_5^2 + 1745944200 f_4^2 f_5) f_7}{1857945600 f_2^{10}} \\
 & + \frac{(1629547920 f_3 f_5 + 1018467450 f_4^2) f_6^2 + 2444321880 f_4 f_5^2 f_6}{1857945600 f_2^{10}}
 \end{aligned} \tag{28b}$$

$$\begin{aligned}
& + \frac{244432188 f_5^4}{1857945600 f_2^{10}} \\
& + \frac{-25865840 f_3^4 f_{10} - 258658400 f_3^3 f_4 f_9}{212336640 f_2^{11}} \\
& - \frac{(465585120 f_3^3 f_5 + 872972100 f_3^2 f_4^2) f_8}{212336640 f_2^{11}} \\
& - \frac{(620780160 f_3^3 f_6 + 2793510720 f_3^2 f_4 f_5 + 1163962800 f_3 f_4^3) f_7}{212336640 f_2^{11}} \\
& - \frac{1629547920 f_3^2 f_4 f_6^2}{212336640 f_2^{11}} \\
& - \frac{(1955457504 f_3^2 f_5^2 + 4888643760 f_3 f_4^2 f_5 + 509233725 f_4^4) f_6}{212336640 f_2^{11}} \\
& - \frac{1955457504 f_3 f_4 f_5^3 - 1222160940 f_4^3 f_5^2}{212336640 f_2^{11}} \\
& + \frac{4759314560 f_3^5 f_9 + 53542288800 f_3^4 f_4 f_8}{5096079360 f_2^{12}} \\
& + \frac{(85667662080 f_3^4 f_5 + 214169155200 f_3^3 f_4^2) f_7}{5096079360 f_2^{12}} \\
& + \frac{49972802880 f_3^4 f_6^2 + (599673634560 f_3^3 f_4 f_5 + 374796021600 f_3^2 f_4^3) f_6}{5096079360 f_2^{12}} \\
& + \frac{119934726912 f_3^3 f_5^3 + 674632838880 f_3^2 f_4^2 f_5^2 + 281097016200 f_3 f_4^4 f_5}{5096079360 f_2^{12}} \\
& + \frac{11712375675 f_4^6}{5096079360 f_2^{12}} \\
& + \frac{-2974571600 f_3^6 f_8 - 35694859200 f_3^5 f_4 f_7}{509607936 f_2^{13}} \\
& - \frac{(49972802880 f_3^5 f_5 + 156165009000 f_3^4 f_4^2) f_6}{509607936 f_2^{13}} \\
& - \frac{187398010800 f_3^4 f_4 f_5^2 - 312330018000 f_3^3 f_4^3 f_5 - 58561878375 f_3^2 f_4^5}{509607936 f_2^{13}}
\end{aligned} \tag{28c}$$

$$\begin{aligned}
& + \frac{(85667662080 f_3^4 f_5 + 214169155200 f_3^3 f_4^2) f_7}{5096079360 f_2^{12}} \\
& + \frac{49972802880 f_3^4 f_6^2 + (599673634560 f_3^3 f_4 f_5 + 374796021600 f_3^2 f_4^3) f_6}{5096079360 f_2^{12}} \\
& + \frac{119934726912 f_3^3 f_5^3 + 674632838880 f_3^2 f_4^2 f_5^2 + 281097016200 f_3 f_4^4 f_5}{5096079360 f_2^{12}} \\
& + \frac{11712375675 f_4^6}{5096079360 f_2^{12}} \\
& + \frac{-2974571600 f_3^6 f_8 - 35694859200 f_3^5 f_4 f_7}{509607936 f_2^{13}} \\
& - \frac{(49972802880 f_3^5 f_5 + 156165009000 f_3^4 f_4^2) f_6}{509607936 f_2^{13}} \\
& - \frac{187398010800 f_3^4 f_4 f_5^2 - 312330018000 f_3^3 f_4^3 f_5 - 58561878375 f_3^2 f_4^5}{509607936 f_2^{13}}
\end{aligned} \tag{28d}$$

$$\begin{aligned}
& + \frac{3399510400 f_3^7 f_7 + 41644002400 f_3^6 f_4 f_6 + 24986401440 f_3^6 f_5^2}{113246208 f_2^{14}} \\
& + \frac{187398010800 f_3^5 f_4^2 f_5}{113246208 f_2^{14}} \\
& + \frac{97603130625 f_3^4 f_4^4}{113246208 f_2^{14}} + \frac{-10782822050 f_3^8 f_6 - 129393864600 f_3^7 f_4 f_5}{84934656 f_2^{15}} \\
& - \frac{188699385875 f_3^6 f_4^3}{84934656 f_2^{15}} \\
& + \frac{1337069934200 f_3^9 f_5 + 7521018379875 f_3^8 f_4^2}{3057647616 f_2^{16}} \\
& - \frac{1838471159525 f_3^{10} f_4}{1528823808 f_2^{17}} + \frac{5849680962125 f_3^{12}}{27518828544 f_2^{18}}
\end{aligned} \quad (28e)$$

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## NOTES

- 1 Contrary to popular belief, an asymptotic expansion is not necessarily a divergent series (Erdelyi, 1956).

## ADDITIONAL INFORMATION

*FIELD:* Mathematics

*ARTICLE TYP:* Review paper

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