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Universitat de Barcelona  
Espanya

BOEDO, F.L.; ESCAYOLA, M.P.; PÉREZ LUJÁN, S.B.; VUJOVICH, G.I.; ARIZA, J.P.;  
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# Geochemistry of Precordillera serpentinites, western Argentina: evidence for multistage hydrothermal alteration and tectonic implications for the Neoproterozoic–early Paleozoic

F.L. BOEDO<sup>1</sup> M.P. ESCAYOLA<sup>1</sup> S.B. PÉREZ LUJÁN<sup>2</sup> G.I. VUJOVICH<sup>1</sup> J.P. ARIZA<sup>3</sup> M. NAIPAUER<sup>1</sup>

<sup>1</sup>Instituto de Estudios Andinos “Don Pablo Groeber” (UBA-CONICET). Departamento de Ciencias Geológicas, Facultad de Ciencias Exactas y Naturales, Universidad de Buenos Aires

Boedo E-mail: florenciabiedo@gmail.com Escayola E-mail: mescayola@gl.fcen.uba.ar

Vujovich E-mail: graciela@gl.fcen.uba.ar Naipauer E-mail: maxinaipauer@gl.fcen.uba.ar

<sup>2</sup>CONICET-Departamento de Geofísica y Astronomía, Facultad de Ciencias Exactas, Físicas y Naturales, Universidad Nacional de San Juan, Argentina

E-mail: sofiap.lujan@unsj-cuim.edu.ar

<sup>3</sup>Instituto Geofísico Sismológico Volponi (UNSJ-CONICET). Universidad Nacional de San Juan, Argentina

E-mail: jpariza@conicet.gov.ar

## ABSTRACT

Serpentinites are a powerful tool to evaluate mantle composition and subsequent alteration processes during their tectonic emplacement. Exposures of this type of rocks can be found in the Argentine Precordillera (Cuyania terrane) and Frontal Cordillera, both located in central-western Argentina, within the Central Andes. In these regions a Neoproterozoic to Devonian mafic-ultramafic belt composed of serpentinites, metabasaltic dikes/sills, pillow lavas (with an Enriched to Normal Mid-Ocean Ridge Basalts (E- to N-MORB) geochemical signature) and mafic granulites crop out, spatially associated with marine metasedimentary rocks. The serpentinite bodies consist of lizardite/chrysotile+brucite+magnetite, with scarce pentlandite and anhedral reddish-brown Cr-spinel (picotite, pleonaste and spinel sensu stricto) as relict magmatic phases. The original peridotites were moderately-depleted harzburgites (ultramafic cumulates) with an intermediate chemical signature between a mid-ocean ridge and an arc-related ophiolite. Whole-rock Rare Earth Elements (REE) patterns of serpentinites exhibit enriched REE patterns ((La/Yb)<sub>CN</sub>=13-59) regarding CI chondrite with positive Eu anomalies. These features are the result of an interaction between hydrothermal fluid and serpentinites, in which moderate temperature (350°–400°C), CO<sub>2</sub>-rich, mildly basic hydrothermal fluid was involved and was responsible for the addition of Ca, Sr and REE to serpentinites. The presence of listvenites (silica-carbonate rocks) in the serpentinite margins allow us to infer another fluid metasomatism, where low-temperatures (<250°C), highly-oxidized, highly-acid fluid lead to the precipitation of silica. The association of these metasomatized serpentinite bodies with neoproterozoic continental margin successions and MORB magmatism at the suture zone of the Cuyania and Chilenia terranes suggests the development of an oceanic basin between them during the Neoproterozoic-early Paleozoic.

## KEYWORDS

Mafic-ultramafic belt. Mid-ocean ridge. Listvenites. Cuyania terrane. Chilenia terrane.

## INTRODUCTION

Mantle rocks tectonically emplaced in orogenic belts provide information about the composition of the mantle and processes that occurred during its formation. Due to the large scale of these processes, they provide information by ultramafic xenoliths, allowing, for example, observation of the relationships between lithologies, evaluation of the mantle heterogeneity scale and analysis of magmatic/metamorphic processes (vein injections, melt-rock reactions, etc.). However, when rock bodies are affected by deformation and metamorphism, their original geodynamic setting is not exactly known and their significance is sometimes a subject of speculation (Bodinier and Godard, 2003). In this case, only relic primary mantle minerals, in particular Cr-spinel, constitute a good petrogenetic indicator that gives information about origin and petrogenesis of mantle peridotite before serpentinization (Pearce *et al.*, 2000; Arai *et al.*, 2011, among others).

Deformational, metamorphic and metasomatic processes can become a fundamental tool for the study of secondary processes. Metasomatic products (metasomatites) are related to the nature of host rocks and the composition of the circulating fluid and are important subjects of mineral research and exploration, since they are valuable indicators for exploration geologists. In particular, the interaction of hydrothermal fluids with ultramafic rocks results in a variety of metasomatic rocks such as quartz-carbonate rocks and serpentine-talc-carbonate rocks. The term “listvenite” was introduced by Rose (1837) to describe silica-carbonate-chromium mica rocks associated with gold-bearing quartz veins. Listvenites are chiefly composed of carbonate (magnesite, dolomite, ankerite), quartz and Cr-mica (mariposite, fuchsite) with disseminated pyrite (Halls and Zhao, 1995) and minor chlorite, and talc. Relic serpentines and Cr-spinel are also found. They are generated by the interaction between intermediate to low temperature hydrothermal/metasomatic alteration of mafic-ultramafic rocks and can be associated with structures (faults, shear zones) or to granitic plutons (Halls and Zhao, 1995; Robinson *et al.*, 2005).

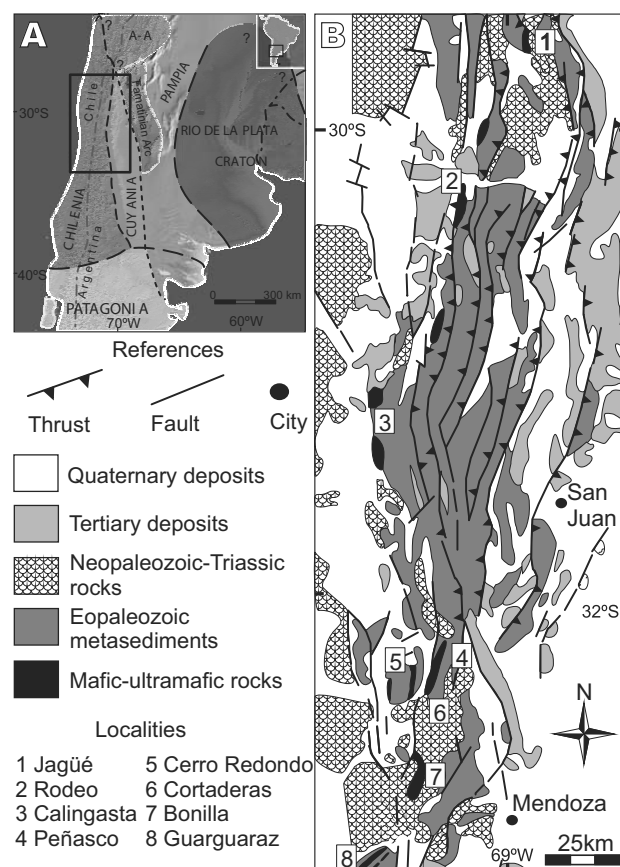
This contribution deals with the alteration products and with relics of primary minerals such as Cr-spinels in an effort to understand the history of the Argentine Precordillera serpentinite bodies, from their tectonic setting of formation to the conditions of metamorphic and metasomatic (listvenitization) processes.

## GEOLOGY OF THE ARGENTINE PRECORDILLERA

### Geological outline

The southwestern margin of the South American plate (28°–33°SL, Fig. 1A) is a natural laboratory where

the tectonic evolution of Gondwana can be observed and analyzed. During early Paleozoic times, the western margin of the continent grew by the accretion of different terranes (Pampia, Cuyania, Chilenia) (Ramos, 2010 and others mentioned therein). Several authors have concluded that the Cuyania terrane (Argentine Precordillera) represents a microcontinent that broke up from southern Laurentia in the early Cambrian and collided against the Gondwana margin during the Late Ordovician (Dickerson, 2012; Thomas *et al.*, 2012). However, the evolution of the western margin of the Cuyania terrane during early Paleozoic times is still a matter of debate. There is a general consensus that during this time span a passive margin setting was developed in association with a shallow ocean basin to the west, which was closed during Middle–Late Devonian times by the collision of the Chilenia terrane against the Gondwana margin (the Cuyania terrane). Different proposed tectonic models (Ramos *et al.*, 1986; Dalla Salda *et al.*, 1992; Loeske, 1993; Davis *et al.*, 1999, 2000, among others) highlight the remaining uncertainties because of the few studies focusing on issues such as the early Paleozoic deformation (Cucchi, 1972; Von Gosen, 1997; Davis *et al.*,



**FIGURE 1.** A) Location of the Chilenia and Cuyania terranes. A-A: Arequipa-Antofalla terrane. Diagram taken from Abre *et al.* (2012). B) Schematic geological map and localities of the Argentine Precordillera mafic-ultramafic belt (after Ramos *et al.*, 2000).

1999; Gerbi *et al.*, 2002), the polarity of the subduction zone between the Chilenia terrane and the Gondwana margin, or the origin and emplacement of the Argentine Precordillera mafic-ultramafic belt, considered as the suture zone between the Chilenia and Cuyania terranes (Haller and Ramos, 1984; Kay *et al.*, 1984; Davis *et al.*, 1999, 2000, among others).

### Geology of the Argentine Precordillera mafic-ultramafic belt

The Argentine Precordillera is a fold and thrust belt located in central-western Argentina, within the subhorizontal subduction segment of the Nazca plate (Central Andes) (Fig. 1A).

In particular, its western sector consists of Ordovician-Devonian metasedimentary rocks of slope and deep marine settings, some of which host Cambrian carbonate olistoliths that belonged to a carbonate platform environment, and siliciclastic rocks from its basement (Thomas and Astini, 2003 and others therein). These associations are spatially related to mafic and ultramafic bodies grouped in the Argentine Precordillera mafic-ultramafic belt (Fig. 1B). This belt crops out discontinuously between 28° and 33°SL and consists of serpentinized ultramafic rocks and massive metagabbros, meta-hyaloclastites, metabasaltic dikes/sills and pillow lavas with greenschist facies metamorphism. The latter show an E-MORB signature and positive  $\epsilon_{\text{Nd}}$  values (+6 to +9.3) that show its oceanic character (Haller and Ramos, 1984; Kay *et al.*, 1984; Cortés and Kay, 1994; Fauqué and Villar, 2003). Mafic granulites retrograded to greenschist metamorphism are also described (Davis *et al.*, 1999; Boedo *et al.*, 2012). Both metaigneous and metasedimentary rocks record a metamorphic event dated by K-Ar and Ar-Ar methods to Middle Devonian age, interpreted as the collision of the Chilenia terrane against the Gondwana margin (Cucchi, 1971; Buggisch *et al.*, 1994; Davis *et al.*, 1999).

Haller and Ramos (1984) considered that the southernmost part of the mafic-ultramafic belt is exposed in the Guarguaraz area, Frontal Cordillera (Fig. 1B). In this area there are outcrops of serpentinized ultramafic bodies, gabbros, basaltic dikes and pillow lavas associated to marbles, mica schists and quartz schists (Villar, 1969, 1970; Bjerg *et al.*, 1990; Gregori and Bjerg, 1997; Gargiulo *et al.*, 2011, among others). U-Pb ages from zircon show a maximum deposition age of  $555 \pm 8$  Ma for these metasediments (Willner *et al.*, 2008), being consistent with findings of probable Vendian-Cambrian cyanobacteria and acritarch fauna and a metabasite with a crystallization age of  $655 \pm 76$  Ma (Sm-Nd whole rock) (López de Azarevich *et al.*, 2009). As in the Precordillera, the Guarguaraz mafic bodies have a N- to E-MORB signature and the whole sequence is strongly metamorphosed and deformed.

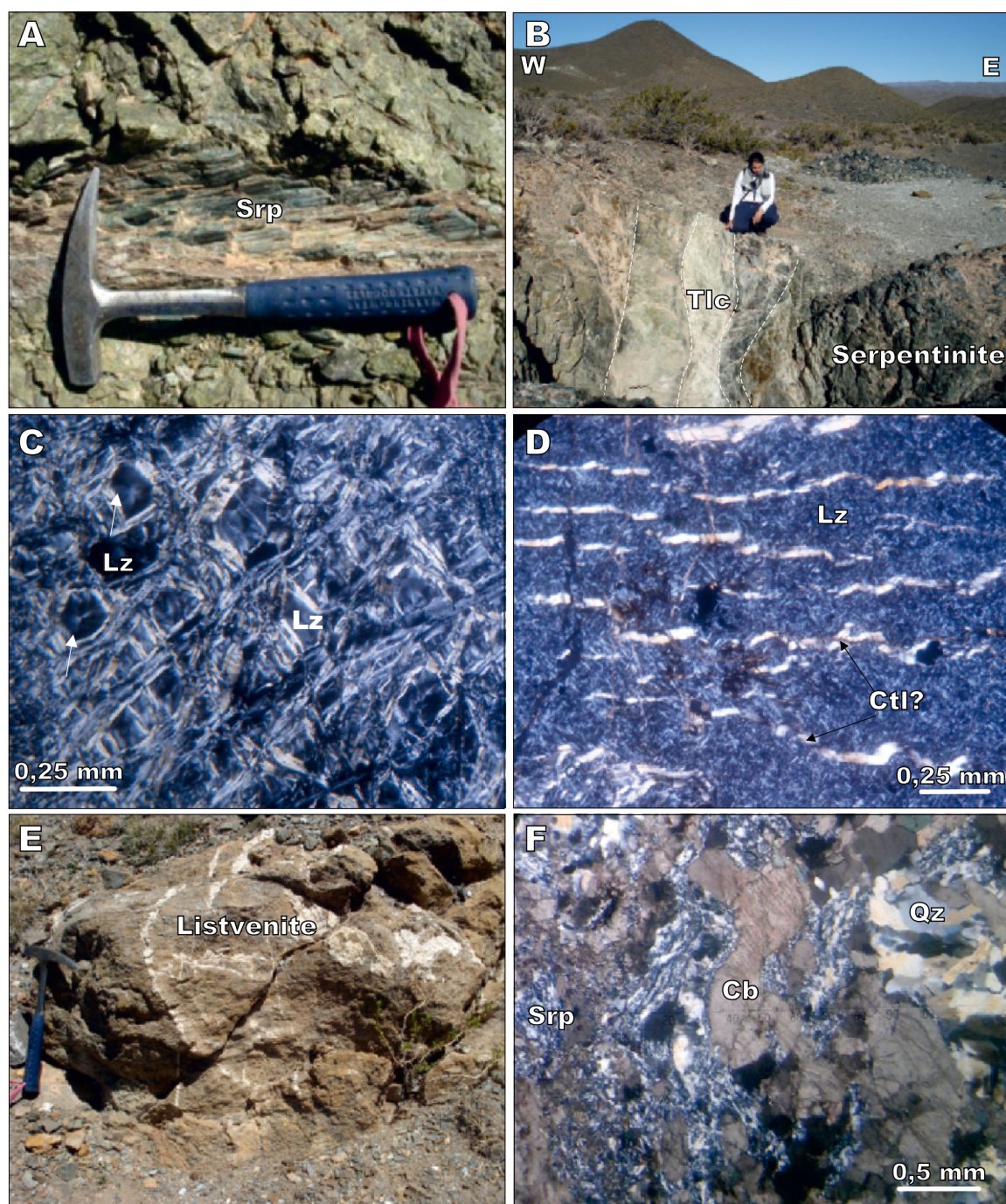
Massonne and Calderón (2008) and Willner *et al.* (2011) estimated anticlockwise P-T paths on metapelites and metabasites yielding high pressure conditions (albite amphibolite facies) followed by a decompression stage, which are compatible with the burial and exhumation of material in a collisional setting. Besides, Willner *et al.* (2011) reported a  $390 \pm 2$  Ma age (Lu-Hf on garnet) interpreted as the time of collision of the Chilenia terrane against the Gondwana margin.

There is some consensus that both localities belong to an almost complete ophiolite sequence formed in a mid-ocean ridge, obducted by the collision of the Chilenia terrane against the Gondwana margin during the Middle-Late Devonian (Haller and Ramos, 1984; Ramos *et al.*, 1984, 1986). Based on U-Pb ages and geochemical data, other authors suggest that mafic and ultramafic bodies would have been generated in different tectonic environments including a mid-ocean ridge (for the pillow lavas, mafic dikes/sills and massive gabbros) and the roots of a magmatic arc (for the mafic granulites) (Davis *et al.*, 1999, 2000). In contrast, Loeske (1993) proposed that the belt belongs to the ocean floor of a back-arc basin based on geochemical and detrital sediment studies.

### Field occurrence and petrographic features of the Peñasco serpentinites

The study zone is located in the Cordón del Peñasco area, at the northernmost part of the province of Mendoza between 32°6'20" and 32°14'19"S and 69°9'15" and 69°3'55"W (Fig. 1B). There deep marine and slope metasedimentary rocks associated with ultramafic rock bodies with advanced to complete serpentinization crop out. In tectonic contact with these bodies are mafic granulites, whose protoliths are layered gabbros since igneous layering is preserved. These bodies are retrograded to greenschist facies metamorphism. In addition, there are basaltic dikes/sills intruding the serpentinites and the metasedimentary successions, assigned to the early Paleozoic and subordinate amygdaloidal metabasites and metahyaloclastites. Both metaigneous and metasedimentary rocks are affected by greenschist facies metamorphism (Harrington, 1971; Boedo *et al.*, 2012; among others).

The serpentinites are widely distributed in the study area, forming north- to northeast-trending west-vergent klippe. They consist of dark to light green, gray-green and black lenticular bodies of variable dimensions (tens to hundreds of meters), with a silky shine and, generally, a massive appearance. Occasionally, stockwork texture is present comprising numerous veinlets and veins a few centimeters to millimeters thick, which are filled with carbonates or serpentines that have grown perpendicularly and/or obliquely to the wall of the veins (Fig. 2A).



**FIGURE 2.** A) Peñasco area. Serpentinite vein filled with fibrous serpentine that has grown obliquely to the wall of the vein. B) Peñasco area. View to the south of a contact zone between a serpentinite (to the east) and a metasedimentary country rock (to the west). Reaction zones of different colors tens of centimeters thick develop. C) Microphotograph of a serpentinite with a mesh texture and hourglass texture in its centre (white arrows). D) Microphotograph of a serpentinite with lizardite interpenetrative texture dissected by subparallel veinlets of possible chrysotile. E) Northern Peñasco area. Rusty orange listvenite body dissected by calcite veins and veinlets of a few centimeters thick. F) Microphotograph of a listvenite (silica-carbonate rock) composed of quartz, carbonate and relic serpentines. Mineral symbols used are those proposed by Whitney and Evans (2010).

The contacts between the serpentinites and their country rock (slates, phyllites and subordinate metacarbonates) are strongly sheared. At these boundaries, concentric/subconcentric reaction zones developed around the ultramafic body, with transitional contacts and thicknesses

from tens of centimeters to a few meters, and variable colors according to mineral content: white (due to the presence of talc), light green (talc+serpentine), bluish gray and black (serpentine+chlorite) (Fig. 2B). Euhedral crystals of pyrite in talc schists are also observed. In many

cases, rusty orange listvenites up to 50 meters wide are associated with serpentinite margins near tectonic contacts with the Peñasco metasedimentary rocks. They consist of quartz, carbonate, chlorite, Fe oxides, relict serpentines and Cr-spinels. As these bodies are associated with major faults whereas intermediate to acidic plutons are relatively far from these rock bodies (up to 100km), we discard an origin related to andesitic-granitic plutons.

Under the microscope serpentinitization is advanced to complete with pseudomorphic/non-pseudomorphic replacement textures, and serpentine veinlets (Figs. 2C–E). Anhedral reddish-brown Cr-spinels and scarce pentlandite and Co-pentlandite constitute the only relict phases. Lizardite mesh with hourglass texture (Fig. 2C) or interpenetrative/cryptocrystalline-interpenetrative texture can be recognized, dissected by veins and veinlets whose thickness varies from <0.1 to 1mm (Fig. 2D). Moreover, blasts of clinochrysotile, magnetite, chlorite, brucite and anhedral crystals of Cr-spinel with chlorite growing at their expense are present (Fig. 2E). In some cases, samples with stockwork textures exhibit up to four generations of veins (Fig. 2E). Occasionally, where serpentinitization is not complete, magnetite and serpentine were recognized along cleavage planes of relict orthopyroxene. Many samples exhibit carbonate patches associated with serpentine and quartz, which manifest the process of listvenitization (Fig. 2F), best developed at the margins of the serpentinite bodies.

## MATERIALS AND METHODS

We collected 11 representative samples of serpentinitized ultramafic rocks from the Argentine Precordillera mafic-ultramafic belt found at the Peñasco, Pozos, Cortaderas and Bonilla localities (Table 1). Eight samples were taken from the Cordón del Peñasco, while the remaining were taken from Pozos, Cortaderas and Bonilla areas (Fig. 1B). We use here the term “serpentinite” to define ultrabasic rock bodies where the serpentinitization process is advanced or complete, such that there are no relics of primary assemblages or original rock textures, and “listvenite” as a silica-carbonate rock product of low-grade metasomatic processes after ultramafic rocks.

The samples were prepared and analyzed following standard procedures. They were milled with a jaw crusher and an agate mortar at the Institute of Geochronology and Isotope Geology (INGEIS) from the Universidad de Buenos Aires-CONICET. Only pieces of the sample which appeared vein-free were used. The samples were analyzed at ACTLABS Laboratories (Canada) where they were prepared and analyzed in a batch system. Each batch contains a method reagent blank, certified reference

material and 17% replicates. Samples were mixed with a flux of lithium metaborate and lithium tetraborate and fused in an induction furnace. The molten melt was poured into a solution of 5% nitric acid containing an internal standard, and mixed until completely dissolved. Samples were run for major oxides and selected trace elements (Ba, Be, Sc, Sr, V, Y, Zr) on an ICP (detection limit: 0.01%, MnO and TiO<sub>2</sub>: 0.001%). Trace elements were measured by Inductively Coupled Plasma Mass Spectrometry (ICP-MS). Results are shown in Table 1, major and trace element concentrations are in weight percent (wt%) and parts per million (ppm), respectively.

The composition of spinel group minerals were analyzed with a JEOL JXA-8230 electron microprobe at the Laboratorio de Microscopía Electrónica y Análisis por Rayos X (LAMARX) of the Universidad Nacional de Córdoba, Argentina. For full quantitative analysis, four wavelength-dispersive spectrometers (WDS) were used. Operating conditions were an acceleration voltage of 15kV for a beam diameter of about 1μm and a beam current of 20nA. Concentrations of Si, Al, Ti, Ca, Mg, Fe, Mn, Na, Cr, V and Zn were determined. Standards were natural minerals: chromite for Mg, Cr and Al; ilmenite for Ti and Fe<sup>2+</sup>; rhodonite for Mn; vanadinite for V; nickeline for Ni; fayalite for Si; anorthoclase for Na; zoisite for Ca and gahnite for Zn. Results are presented in Table 2. Fe<sup>3+</sup> contents were estimated by the method proposed by Droop (1987).

In order to determine the type of serpentine present, X-ray diffraction studies were performed at Universidad Nacional de Córdoba with a diffractometer Rigaku D-Max IIC. Operating conditions were: 5–70° 2θ, velocity 3°/minute, step 0.02, radiation Cu (Kα=1.5405Å). Software employed was MDI Jade 7.

## RESULTS AND DISCUSSION

### Composition of relict Cr-spinel

Cr-spinels can be recognized by their intercumulus relic disposition (Fig. 3A) or by their holly leaf shapes (Fig. 3B). Their rims are usually altered to ferritchromite, magnetite or Cr-rich magnetite (Fig. 3C), but occasionally the whole crystal is altered showing spongy appearance (Fig. 3D).

Fresh crystals are unzoned (Fig. 3C) and exhibit an homogeneous composition, with Cr# (=Cr/(Cr+Al) atomic ratio) between 0.35 and 0.50 and Mg# (=Mg/(Mg+Fe<sup>2+</sup>) atomic ratio) between 0.56–0.72. MnO, V<sub>2</sub>O<sub>3</sub> and ZnO do not exceed 0.4%. Chromites from sample LJ 22-12 show high TiO<sub>2</sub> contents between 0.4–0.5% and restricted Al<sub>2</sub>O<sub>3</sub>

**TABLE 1.** Chemical data set of the Peñasco serpentinites. Major and trace element contents are given in wt% and ppm, respectively

Locality	Northern Cordón del Peñasco			Central Cordón del Peñasco					Pozos area	Bonilla area	Cortaderas area
Sample	EN 1-12	EN 3-12	EN 12-12	LJ 21-12	QM 7-08	QM 8-08	QM 28-08	QM 38-08	QP 7	SE 15A	M1
SiO <sub>2</sub>	39,83	38,63	40,17	38,37	39,22	34,12	40,28	40,74	38,27	40,40	39,88
TiO <sub>2</sub>	b.d.l. <sup>1</sup>	0,04	0,02	0,01	0,01	0,01	0,06	0,01	0,02	0,07	0,02
Al <sub>2</sub> O <sub>3</sub>	0,72	1,76	1,47	1,86	0,65	0,62	1,88	0,35	1,95	2,30	1,08
Fe <sub>2</sub> O <sub>3</sub> T	5,64	6,01	5,51	6,63	6,78	6,28	5,42	6,95	6,32	6,19	5,98
MnO	0,05	0,11	0,08	0,09	0,09	0,10	0,10	0,08	0,12	0,09	0,08
MgO	36,21	37,14	37,40	36,67	35,22	30,78	36,31	36,65	36,14	36,73	36,29
CaO	1,18	0,11	0,34	0,08	2,94	10,60	0,77	0,12	0,19	0,06	1,46
Na <sub>2</sub> O	0,01	b.d.l	0,02	b.d.l	0,02	0,03	0,02	0,01	b.d.l	b.d.l	b.d.l
K <sub>2</sub> O	b.d.l	b.d.l	b.d.l	b.d.l	0,01	0,01	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l
P <sub>2</sub> O <sub>5</sub>	b.d.l	0,02	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l
LOI	15,53	13,38	13,67	13,16	14,44	17,71	13,98	12,10	13,56	12,14	12,41
Total	99,16	97,20	98,68	96,87	99,37	100,30	98,81	97,00	96,56	97,97	97,20
La	8,25	9,00	3,80	3,89	0,17	0,19	3,18	3,81	8,73	3,50	6,38
Ce	11,80	14,30	4,37	4,20	0,42	0,41	4,72	5,32	12,80	4,49	10,40
Pr	1,48	1,71	0,53	0,77	0,04	0,05	0,62	2,05	1,40	0,67	1,21
Nd	4,38	5,73	1,60	2,16	0,13	0,14	1,72	2,13	3,70	2,16	3,80
Pm	-	-	-	-	-	-	-	-	-	-	-
Sm	0,72	0,94	0,30	0,49	0,01	0,01	0,32	0,38	0,40	0,31	0,59
Eu	0,21	0,30	0,15	0,13	0,02	0,05	0,15	0,36	0,71	0,33	0,14
Gd	0,30	0,58	0,14	0,18	0,03	0,02	0,21	0,17	0,21	0,25	0,31
Tb	0,04	0,09	0,03	0,03	b.d.l	b.d.l	0,04	0,03	0,04	0,04	0,04
Dy	0,18	0,45	0,18	0,18	0,04	0,03	0,22	0,15	0,23	0,23	0,18
Ho	0,03	0,08	0,04	0,04	b.d.l	b.d.l	0,04	0,03	0,05	0,05	0,03
Er	0,09	0,22	0,13	0,11	0,03	0,01	0,14	0,09	0,15	0,15	0,09
Tm	0,01	0,03	0,02	0,02	b.d.l	b.d.l	0,02	0,01	0,03	0,03	0,02
Yb	0,10	0,20	0,14	0,12	0,03	0,01	0,16	0,10	0,20	0,19	0,13
Lu	0,02	0,03	0,02	0,02	0	b.d.l	0,03	0,02	0,04	0,04	0,03
Sr	33	8	8	4	24	67	16	3	4	3	20
Ba	7	6	7	b.d.l	7	9	8	3	6	b.d.l	b.d.l
Cs	0,50	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l
Rb	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l
U	0,99	0,09	0,27	0,16	1	2	0,14	1,39	0,23	0,05	b.d.l
Th	0,49	0,65	0,16	0,19	0	b.d.l	0,23	0,23	0,50	0,19	0,42
Hf	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l
Nb	b.d.l	0,6	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l
Ta	0,19	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	0,02	b.d.l	b.d.l	b.d.l	b.d.l
Sc	6	9	8	9	6	4	10	3	11	11	6
Cr	2280	2480	2310	2200	2460	2340	2650	2210	2670	2200	2070
Ni	2080	2130	2160	2040	2400	1690	1970	1990	2060	1830	1970
Co	99	107	102	104	105	75	88	93	103	91	105
Pb	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l
Nb	b.d.l	0,60	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l	b.d.l
Zr	b.d.l	2	b.d.l	b.d.l	2	b.d.l	2	b.d.l	b.d.l	1	b.d.l
V	22	49	40	39	25	25	53	25	57	62	37
Y	b.d.l	2,00	0,90	1,00	b.d.l	b.d.l	1,00	b.d.l	1,20	1,00	0,70

<sup>1</sup>b.d.l: below detection limit

**TABLE 2.** Chemical data set of the Cr-spinel group minerals from the Peñasco serpentinites. Element contents are given in wt%

Sample	EN 16-12										LJ 21-12				LJ 22-12				
SiO <sub>2</sub>	-	-	-	-	-	-	-	-	-	-	-	-	0,12	-	-	-	-	-	-
TiO <sub>2</sub>	0,11	0,12	0,07	0,13	0,10	0,09	0,06	0,09	0,09	0,09	0,05	0,04	0,05	0,08	0,48	0,47	0,45	0,41	0,51
Al <sub>2</sub> O <sub>3</sub>	32,71	30,59	33,49	32,39	37,20	37,65	36,83	36,47	35,84	34,04	36,33	33,99	27,39	28,79	30,73	30,84	30,80	29,91	29,87
Cr <sub>2</sub> O <sub>3</sub>	35,12	38,16	35,05	33,73	30,02	29,98	30,74	31,24	32,34	34,60	30,99	33,78	36,91	36,45	33,46	33,16	33,61	33,04	33,44
NiO	0,08	-	0,16	0,12	0,16	0,10	0,14	0,16	0,08	0,14	0,18	0,13	0,08	0,03	0,12	0,10	0,05	0,15	0,12
V <sub>2</sub> O <sub>3</sub>	0,05	0,25	0,07	0,13	0,22	0,16	0,16	0,21	0,14	0,18	0,13	0,16	0,20	0,18	0,16	0,15	0,21	0,29	0,22
FeO	15,48	15,89	15,71	16,23	15,20	14,88	15,26	14,89	14,03	15,09	14,85	15,76	18,63	18,15	18,54	18,52	19,30	18,59	19,38
MgO	15,53	15,40	16,11	14,80	16,29	16,53	16,44	16,37	16,68	16,77	17,29	16,44	13,54	14,09	14,93	14,95	15,32	14,57	14,69
MnO	0,27	0,23	0,24	0,27	0,23	0,23	0,18	0,23	0,24	0,27	0,33	0,24	0,35	0,27	0,27	0,23	0,27	0,27	0,27
CaO	-	-	-	-	-	-	0,03	-	-	-	-	-	0,02	-	-	-	-	-	-
ZnO	0,25	0,13	0,26	0,10	0,19	0,23	0,20	0,21	0,13	0,22	0,28	0,10	0,28	0,19	0,18	0,17	0,12	0,10	0,16
Na <sub>2</sub> O	0,01	-	0,02	0,01	0,02	-	0,03	-	-	-	-	-	-	-	-	-	-	-	-
Total	99,43	100,57	101,05	97,81	99,52	99,66	99,96	99,71	99,39	101,23	100,29	100,46	97,57	98,03	98,87	98,58	100,13	97,32	98,66
Si	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-	-
Ti	0,002	0,003	0,002	0,003	0,002	0,002	0,001	0,002	0,002	0,002	0,001	0,001	0,001	0,002	0,011	0,011	0,010	0,010	0,012
Al	1,135	1,060	1,141	1,144	1,261	1,271	1,246	1,237	1,219	1,152	1,225	1,160	1,000	1,039	1,104	1,109	1,094	1,094	1,081
Cr	0,817	0,887	0,801	0,799	0,683	0,679	0,698	0,711	0,738	0,785	0,701	0,773	0,904	0,882	0,806	0,800	0,801	0,811	0,812
Ni	0,002	-	0,004	0,003	0,004	0,002	0,003	0,004	0,002	0,003	0,004	0,003	0,002	0,001	0,003	0,002	0,001	0,004	0,003
V	0,001	0,006	0,002	0,003	0,005	0,004	0,004	0,005	0,003	0,004	0,003	0,004	0,005	0,004	0,004	0,004	0,005	0,007	0,005
Fe <sup>2+</sup>	0,381	0,391	0,380	0,407	0,366	0,357	0,366	0,359	0,338	0,362	0,355	0,382	0,483	0,465	0,425	0,425	0,438	0,434	0,448
Mg	0,681	0,675	0,694	0,661	0,699	0,706	0,703	0,703	0,717	0,717	0,737	0,710	0,626	0,643	0,678	0,680	0,688	0,674	0,672
Mn	0,007	0,006	0,006	0,007	0,006	0,006	0,004	0,006	0,006	0,007	0,008	0,006	0,009	0,007	0,007	0,006	0,007	0,007	0,007
Ca	-	-	-	-	-	-	0,001	-	-	-	-	-	0,001	-	-	-	-	-	-
Zn	0,006	0,003	0,006	0,002	0,004	0,005	0,004	0,005	0,003	0,005	0,006	0,002	0,006	0,004	0,004	0,004	0,003	0,002	0,004
Na	0,001	-	0,001	0,001	0,001	-	0,002	-	-	-	-	-	-	-	-	-	-	-	-
Fe <sup>3+</sup>	0,07	0,06	0,08	0,07	0,07	0,07	0,08	0,07	0,06	0,08	0,10	0,09	0,11	0,11	0,10	0,10	0,11	0,10	0,11
Cr#	0,42	0,46	0,41	0,41	0,35	0,35	0,36	0,36	0,38	0,41	0,36	0,40	0,47	0,46	0,42	0,42	0,42	0,43	0,43
Mg#	0,64	0,63	0,65	0,62	0,66	0,66	0,66	0,66	0,68	0,66	0,67	0,65	0,56	0,58	0,61	0,62	0,61	0,61	0,60

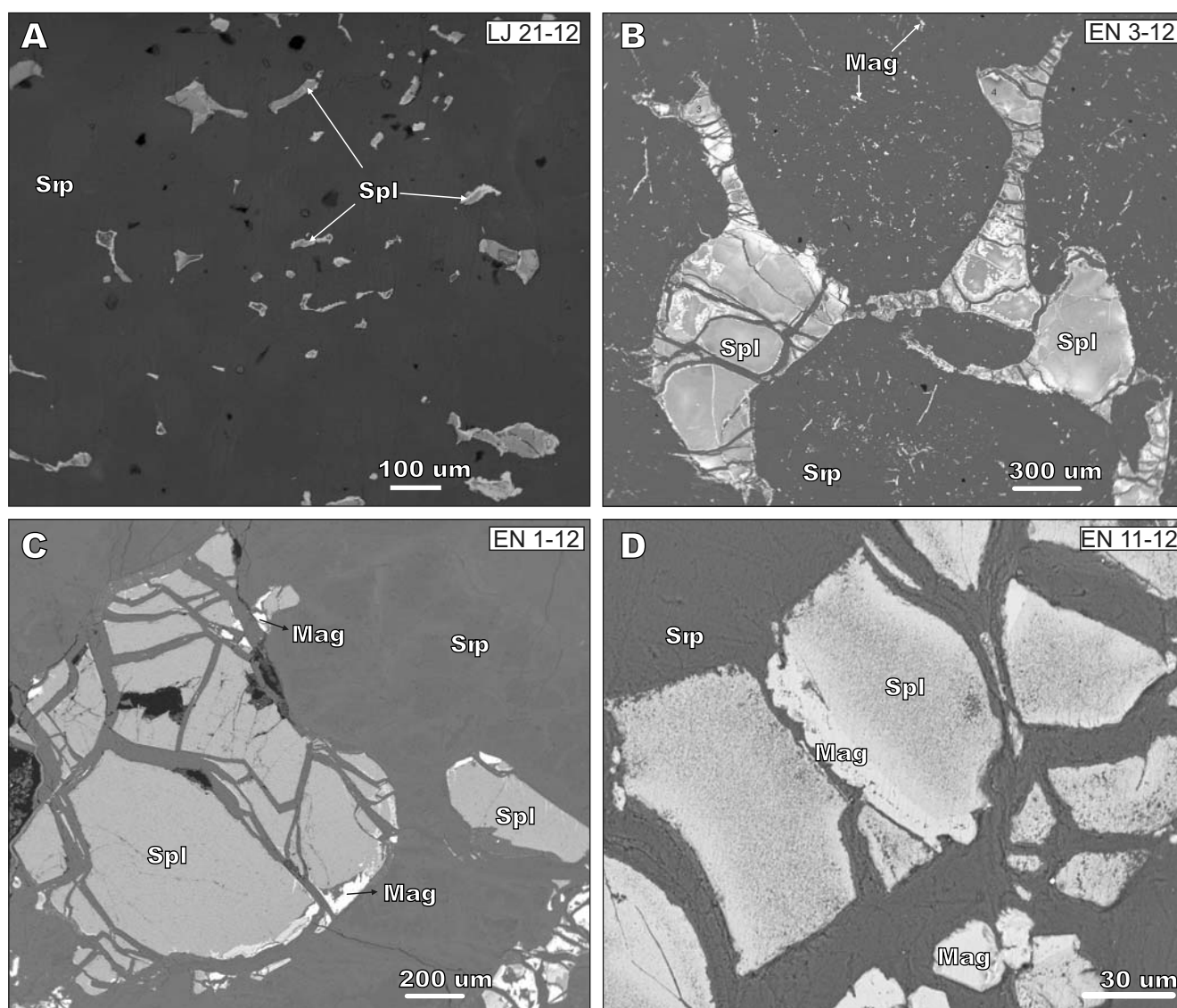
concentrations, whereas chromites from the remaining samples (EN 1-12, 16-12 and LJ 21-12) are poorer in TiO<sub>2</sub> (up to 0.1%) and richer in Al<sub>2</sub>O<sub>3</sub> and Cr<sub>2</sub>O<sub>3</sub> (Fig. 4A–C). Chromites from sample EN 1-12 exhibit the highest Cr<sub>2</sub>O<sub>3</sub> and MgO values (Fig. 4D).

Unusual compositions can be attributed to metasomatism (Mukherjee *et al.*, 2010; Singh and Singh, 2013, among others) and can be separated in two groups: 1) MnO–ZnO rich and MgO–Al<sub>2</sub>O<sub>3</sub> poor crystals, and 2) MnO rich and Al<sub>2</sub>O<sub>3</sub>–Cr<sub>2</sub>O<sub>3</sub> poor crystals. The first group shows MnO contents up to 9.6% and ZnO concentrations that occasionally exceed 1.5%. Al<sub>2</sub>O<sub>3</sub> does not exceed 3% and MgO contents reach 5%. The second group exhibits MnO contents of 3.5–4.7%, with Al<sub>2</sub>O<sub>3</sub> and Cr<sub>2</sub>O<sub>3</sub> concentrations up to 6.7% and 17.8%, respectively. The presence of frequent silicate microinclusions are shown by high SiO<sub>2</sub> contents (up to 20%, analyses not included) in Cr-spinel analyses.

Spinel is classified as picotites and pleonaste and, to a lesser extent, spinel *sensu stricto* (Fig. 5) according to the chemical variation of spinels (solid solution spinel-hercynite-chromite-magnesiochromite-magnesioferrite-magnetite).

Cr-rich spinels are good petrogenetic indicators of mafic-ultramafic rocks. In particular, Cr# rarely exceed

0.6 in Mid-Ocean-Ridge (MOR) rocks (0.2–0.6 according to Haggerty, 1991), whereas arc rocks can reach values of 0.9, and intraplate settings influenced by mantle plumes have >0.7 (Haggerty, 1991). It is broadly accepted that Cr# values >0.6 reflect high degree of melting (Dick and Bullen, 1984; Arai *et al.*, 2011), so mid-ocean-ridge ultramafic rocks will exhibit Cr# between 0.2 and 0.6 (Coish and Gardner, 2004; Deschamps *et al.*, 2013). Cr# and TiO<sub>2</sub> spinel contents suggest that the majority of the Peñasco serpentinites experienced a significant degree of partial melting (>20%) (Fig. 6A). Cr-spinels from sample LJ 22-12 lie in a melt-mantle interaction trend similar to the one described by Pearce *et al.* (2000) for the South Sandwich trench-fracture zone intersection peridotites. This sample could have originated from an interaction between a MORB-like melt and a mantle that had experienced relatively little partial melting (up to 10%). From Cr# versus Mg# diagram, there is no obvious discrimination since samples plot in the superposition zone of Forearc, Backarc and Abyssal peridotite settings (Fig. 6B). In accordance, on the TiO<sub>2</sub> versus Al<sub>2</sub>O<sub>3</sub> diagram (Kamenetsky *et al.*, 2001), the Peñasco Cr-spinels plot within the fields of the MORB and suprasubduction zone (SSZ) peridotites (Fig. 6C), whereas on the Cr# versus Fe<sup>2+</sup>/(Mg<sup>2+</sup>+Fe<sup>2+</sup>) diagram, samples plot within the array of ocean floor peridotites (Fig. 6D) and follow the Ca–Al trend proposed by Barnes and Roeder (2001). This trend shows a widely variable Cr# at generally low Fe<sup>2+</sup>/(Mg<sup>2+</sup>+Fe<sup>2+</sup>). It



**FIGURE 3.** A) Back scattered electron (BSE) image of a serpentinite with relict intercumulus Cr-spinels. B) BSE image of a serpentinite with metasomatized Cr-spinels showing holly leaf shapes. C) BSE image of chemically homogeneous Cr-spinels with a thin rim of magnetite. D) BSE image of a Cr-spinel almost completely altered to magnetite. Mineral symbols used are those proposed by Whitney and Evans (2010).

was first described by Irvine (1967), who explained that it corresponds to spinels equilibrating with olivine at constant composition and temperature. It is particularly evident in the various mantle and lower-crustal samples such as xenoliths, ophiolites and ocean-floor peridotites (Barnes and Roeder, 2001). From Cr-spinel chemistry, the Peñasco serpentinites exhibit an intermediate signature between MOR and arc-related settings.

#### Geochemical aspects of the Peñasco serpentinites

The whole rock analyses show that the Peñasco samples have  $\text{SiO}_2$  concentrations that vary over a small range (45.8–47.9wt%), with  $\text{Al}_2\text{O}_3$  between 0.4 and 2.6wt% and Mg# varies from 0.93 to 0.94 (MgO varying between 42.8

and 44.3wt%). Samples reach  $\text{TiO}_2$  concentrations up to 0.07wt% and CaO values do not exceed 2.9wt%. Some samples analyzed show addition of CaO up to 10wt% (Table 1), which evidence the carbonatization. Loss on ignition (LOI) values are high (12.1–17.7%), which could be considered as a parameter of the advanced to complete serpentinization process. However, some authors suggest that, among serpentines, carbonates and other hydrous minerals such as brucite could influence these values (Deschamps *et al.*, 2013). Surely, the latter is the case for several of the studied samples since pure serpentinite has 13wt%  $\text{H}_2\text{O}$ .

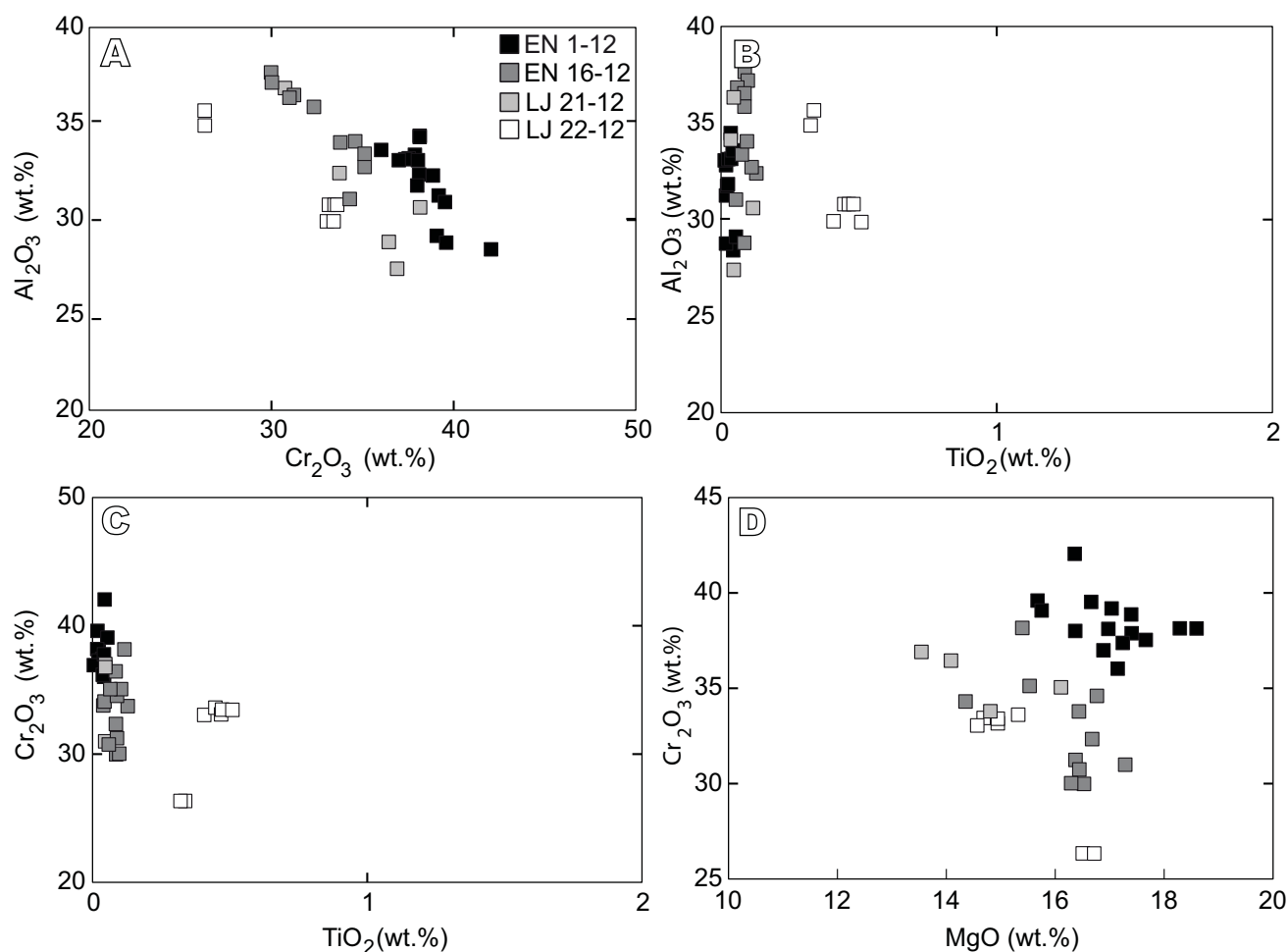
Cr and Ni contents are between 2070–2650ppm and 1690–2400ppm, respectively. Co concentrations range

from 75 to 107ppm, whereas V contents vary between 22 and 62ppm. We interpret a cumulate origin for the Peñasco serpentinites since these values are similar to those informed by Coleman (1977) for ophiolite ultramafic cumulates. In general, when normalized to chondrite CI (Sun and McDonough, 1989), serpentinites show enriched REE patterns up to 40 times higher than the chondritic values, with an evident relative Light Rare Earth Elements (LREE) enrichment over Heavy Rare Earth Elements (HREE) ( $(La/Yb)_{CN}=13-59$ ), and marked positive Eu anomalies ( $Eu/Eu^*$  up to 7.5, being  $Eu^*=(Sm_N \times Gd_N)^{1/2}$  –Taylor and McLennan, 1985) (Fig. 7A). In particular, there is a LREE fractionation with  $(La/Sm)_{CN}$  values between 5.1 and 14, while HREE have subhorizontal to slightly positive slopes ( $(Sm/Yb)_{CN}$ : 1.8–8), which indicate that protoliths were at a mantle depth out of the garnet stability field.

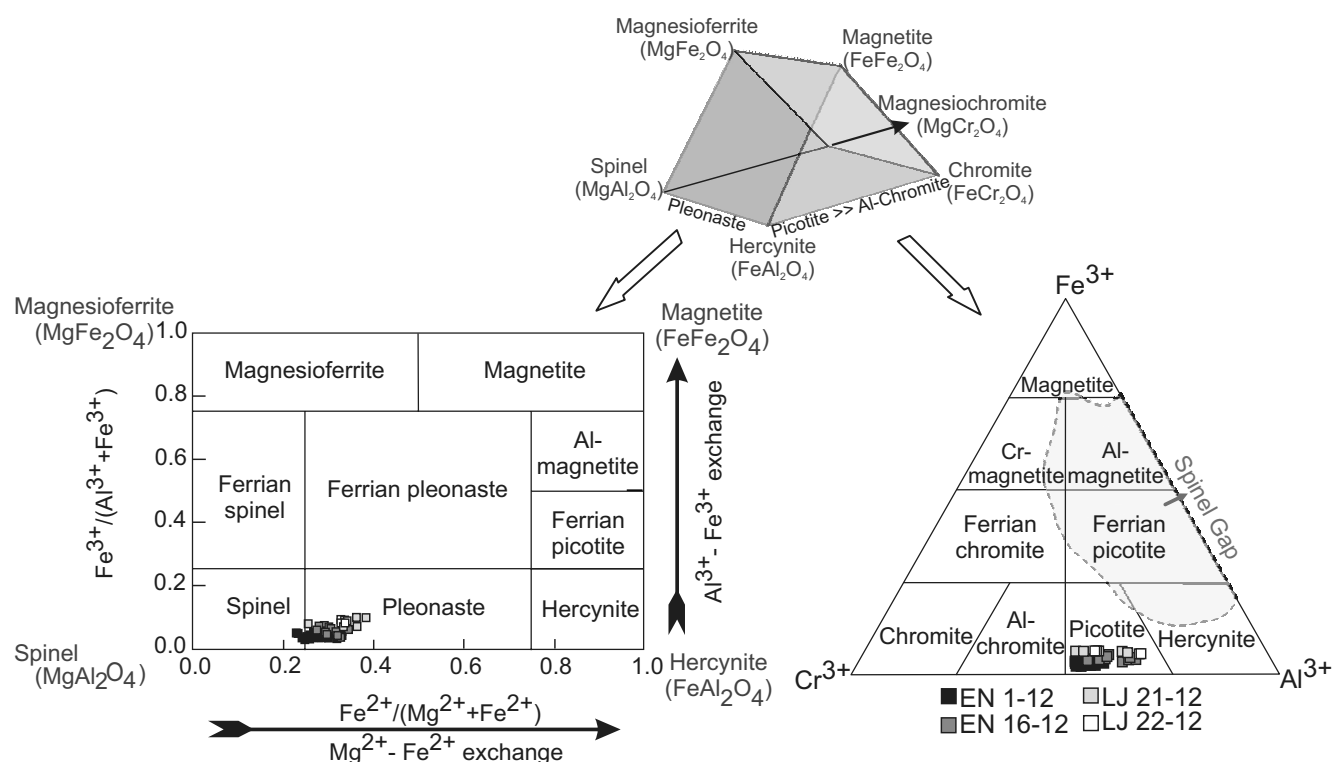
Serpentinites exhibit Th and U (and occasionally Cs) enrichment and negative anomaly on HFSE such as Zr and Ti (Fig. 7B). U enrichments and Cs erratic values can

be explained by weathering processes occurring near the seafloor (Kodolányi *et al.*, 2012 and others mentioned therein).

Neither the high REE contents nor the positive Eu anomalies (Figs. 7A) could have been inherited from the original ultramafic mineral assemblage, since olivine and pyroxene are not important REE carriers and there is no evidence of relict plagioclase or mantle refertilization. Some samples also exhibit moderate Sr concentrations, which are affected by serpentinization or carbonatization. In carbonate-bearing serpentinites, carbonates are the most important Sr carriers (between 1 and 2 orders of magnitude higher than in serpentine), while in carbonate-free serpentinites, serpentines are the main Sr carriers (Kodolányi *et al.*, 2012). In the examined serpentinites the amount of carbonate increase toward the margins, where listvenites can be found. It is noteworthy that despite the fact that all analyzed serpentinites show the same REE enriched patterns, not all of them contain carbonates in hand specimen or in thin section. In some cases, the only



**FIGURE 4.** A)  $Al_2O_3$  versus  $Cr_2O_3$  diagram showing trends of analyzed Cr-spinels from different serpentinite samples. B)  $Al_2O_3$  versus  $TiO_2$  diagram. C)  $Cr_2O_3$  versus  $TiO_2$  diagram. D)  $Cr_2O_3$  versus  $MgO$  diagram. Oxides are expressed in weight percent (wt%).



**FIGURE 5.** Chemical classification diagrams for the spinel group minerals of the Peñasco area. Spinel prism for the multi-component system: spinel-hercynite-chromite-magnesiochromite-magnesioferrite-magnetite. Binary classification diagram considering the  $\text{Mg}^{2+}$ - $\text{Fe}^{2+}$  and  $\text{Al}^{3+}$ - $\text{Fe}^{3+}$  exchanges (left). Triangular classification diagram  $\text{Cr}^{3+}$ - $\text{Fe}^{3+}$ - $\text{Al}^{3+}$ . Spinel gap field from Barnes and Roeder (2001) (right). Diagrams taken from Gargiulo *et al.* (2013).

evidence for metasomatism are chemically altered relict Cr-spinels (discussed above). Hence, the metasomatism in some serpentinites could be cryptic as well.

#### Hydrothermal alteration and listvenitization

From their original setting through their tectonic emplacement in metasediments, the Peñasco serpentinite bodies exhibit a complex history of alteration and metasomatism. At first, the ultramafic bodies were affected by serpentinization under low pressure and temperature conditions (hydrothermal alteration of ocean floor) mainly evidenced by the presence of lizardite and chrysotile forming pseudomorphic textures after olivine (mesh and hourglass textures). The recognized assemblage lizardite+chrysotile+magnetite±brucite±chlorite indicates that the maximum temperature reached during ocean-floor alteration was  $\leq 300^\circ\text{C}$  (Evans, 1977, 2004).

Ideally, serpentinites can be described as a series of minerals with compositions essentially on a tie-line between  $\text{Mg}(\text{OH})_2$  (brucite) and  $\text{Mg}_3\text{Si}_4\text{O}_{10}(\text{OH})_2$  (talc) in the system  $\text{MgO}$ - $\text{SiO}_2$ - $\text{H}_2\text{O}$  (MSH) (Mellini *et al.*, 1987). At low temperature conditions ( $< 300^\circ\text{C}$ ), serpentinites consist mainly of lizardite and chrysotile (Evans, 1977, 2004). Both minerals are replaced by antigorite during

prograde metamorphism. Antigorite stability field ranges from  $> 300^\circ\text{C}$  to  $600$ – $640^\circ\text{C}$ , which constitute a temperature range of  $300^\circ\text{C}$  or more (Wunder *et al.*, 2001; Bromiley and Pawley, 2003).

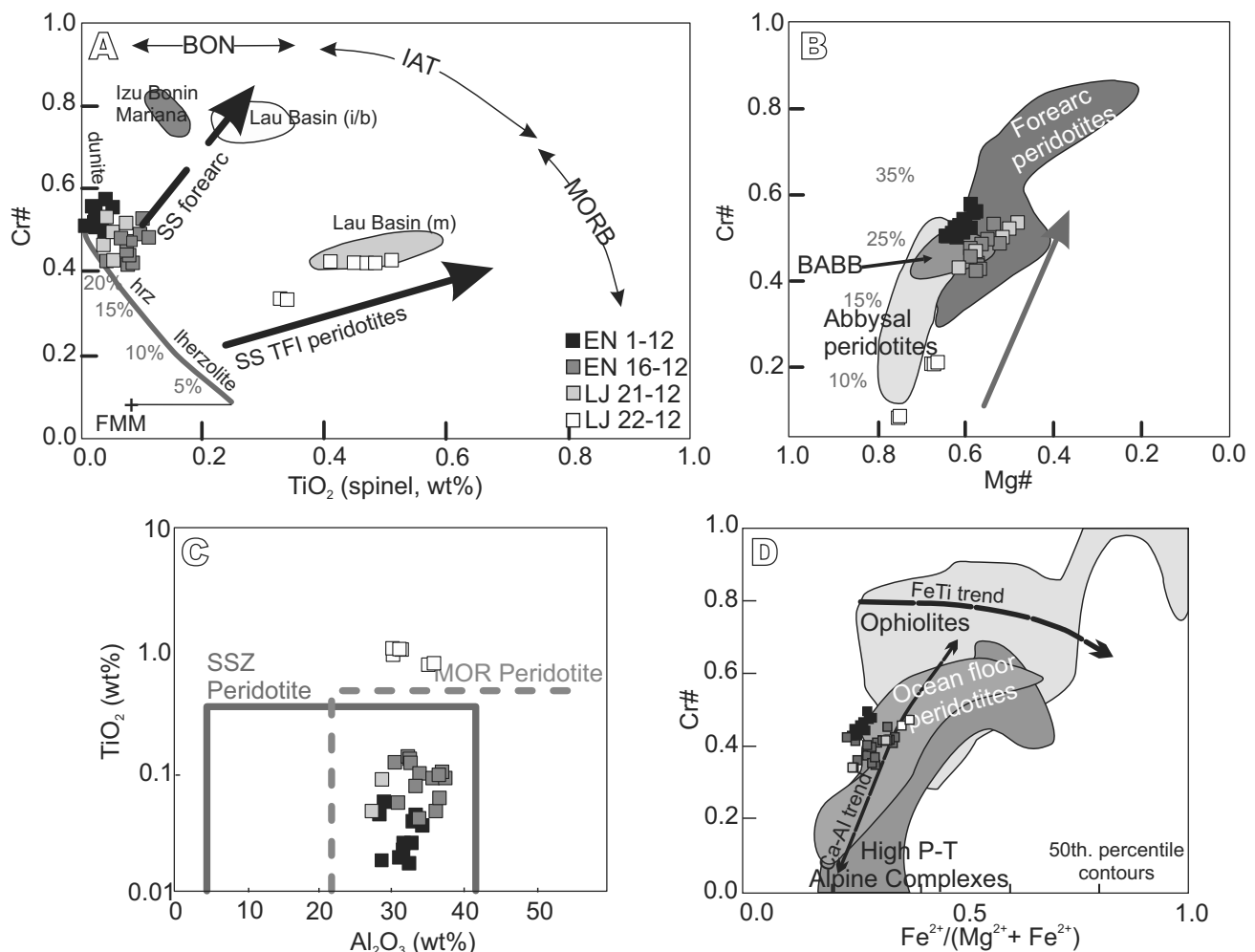
From petrographic and geochemical features already described and parageneses defined by Evans (1977, 2004), Winkler (1978) and Spear (1995), the Peñasco serpentinites are considered in the  $\text{MgO}$ - $\text{SiO}_2$ - $\text{H}_2\text{O}$ - $\text{Al}_2\text{O}_3$ - $\text{FeO}$  (MSHAF) system and are characterized by high initial Fo/En and affected by greenschist facies metamorphism. Occasionally, chlorite (aluminous phase) occurs at Cr-spinel margins and might have gained aluminum during its magnetization.

We do not reject the presence of antigorite in samples not analyzed by X-ray diffraction. From microprobe analyses, we have detected some serpentine patches with  $\text{SiO}_2$  contents similar to those of antigorite crystals compiled by Deer *et al.* (1962). If antigorite were formed during low-grade metamorphism, the assemblage antigorite+brucite would suggest temperatures up to  $400^\circ\text{C}$ , in accordance with the chemistry of metasomatic fluids already discussed. We also infer that the fluids involved in the metamorphic event would have had low  $\text{XCO}_2$  because if  $\text{XCO}_2$  were  $> 10\%$ , serpentines would become unstable (Winkler, 1978).

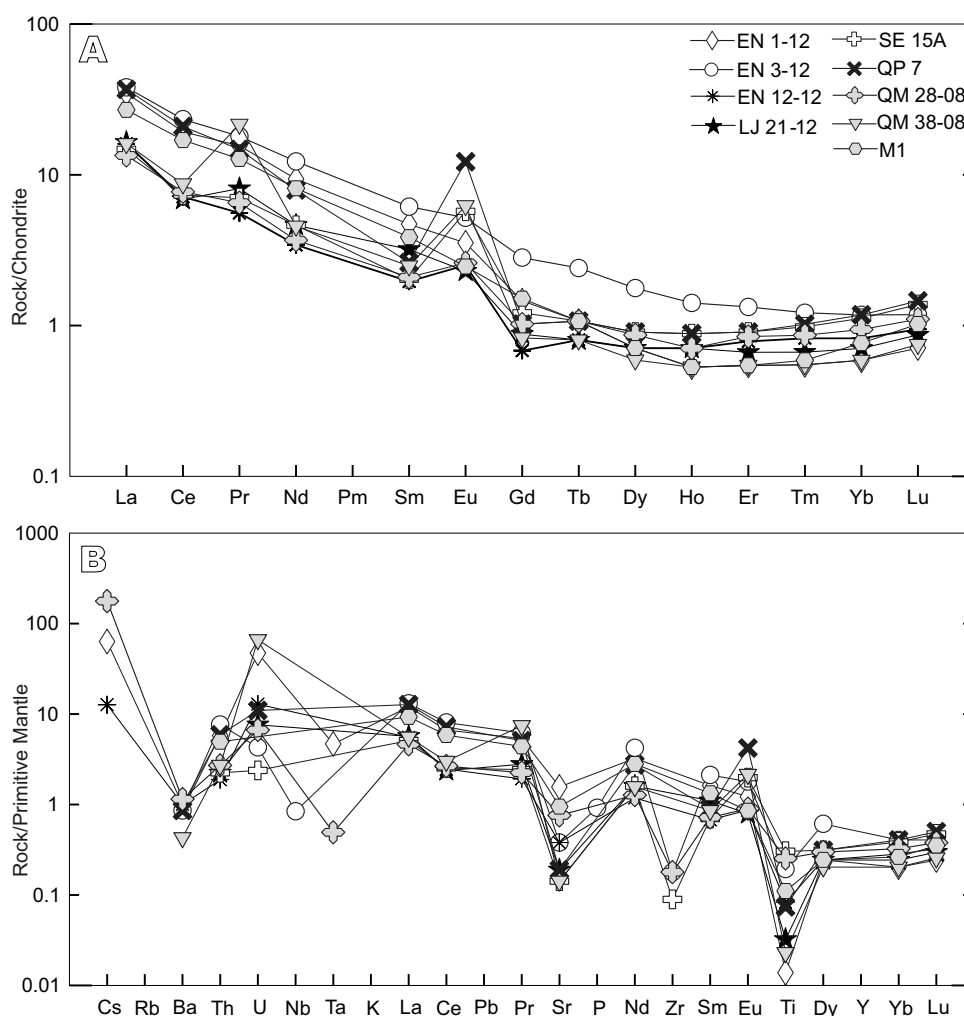
REE mobilization, Eu positive anomalies, Ca and Sr addition and the presence of carbonate patches can be related to a hydrothermal fluid/serpentinite interaction, with high fluid/rock ratios ( $>100$ ), at temperatures below  $400^{\circ}\text{C}$ , as has been interpreted by Menzies *et al.* (1993), Allen and Seyfried (2005) and Paulick *et al.* (2006). Fluids rich in  $\text{CO}_2$  and complexing ions such as  $\text{F}^-$ ,  $\text{Cl}^-$ ,  $\text{PO}_4^{3-}$ ,  $\text{CO}_3^{2-}$ ,  $\text{SO}_4^{2-}$ , among others, coupled with low pH promote REE mobilization (Wood, 1990). During cooling, calcite is known to preferentially incorporate LREE due to the fact that the difference in ionic radius is smaller between Ca and LREE than between Ca and HREE. The marked positive Eu anomalies observed could be related to the

presence of Eu as  $\text{Eu}^{2+}$ , which is easily exchanged with Ca. The presence of abundant carbonate in some samples and the LREE fractionation relative to the HREE suggest an important contribution of  $\text{CO}_3^{2-}$  ions, which fractionate LREE relative to HREE, implying mildly basic (higher pH) conditions for REE-carbonate complexation (Tsikouras *et al.*, 2006).

The chemistry of the fluid described above is consistent with the absence of silica in serpentinites, which indicates higher pH and moderate to high temperature fluid ( $350^{\circ}\text{--}400^{\circ}\text{C}$ ), that holds silica in solution and favors carbonate precipitation (Uçurum, 2000). Overall, the



**FIGURE 6.** Tectonic discrimination diagrams for spinel group minerals from the Peñasco serpentinites. A)  $\text{Cr}^{\#}_{\text{spl}}$  versus  $\text{TiO}_2$  (spl, wt%) diagram that discriminates between ultramafic rocks that were affected only by partial melting and those which also experienced melt-mantle interaction. Analyzed Cr-spinels show that samples EN 1-12, EN 16-12 and LJ 21-12 experienced significant partial melting ( $>20\%$ ), whereas sample LJ 22-12 could have been originated by an interaction between a MORB-like melt and a mantle that had experienced relatively little partial melting ( $10\text{--}15\%$ ), since its trend is similar to that described by Pearce *et al.* (2000) for the South Sandwich Trench-Fracture Intersection zone peridotites. From Pearce *et al.* (2000). B)  $\text{Cr}^{\#}_{\text{spl}}$  versus  $\text{Mg}^{\#}_{\text{spl}}$  diagram, where forearc, backarc and abyssal peridotites are represented. Except for LJ 22-12, analyzed Cr-spinels plot in the superposition zone of the aforementioned fields. BABB: Back-Arc Basin Basalts. C)  $\text{TiO}_2$  (spl, wt%) versus  $\text{Al}_2\text{O}_3$  (spl, wt%) diagram (from Kamenetsky *et al.*, 2001).  $\text{Al}_2\text{O}_3$  values  $>20\%$  imply high degrees of mantle melting and are typical of spinels from mid-ocean ridge peridotites. Cr-spinels from the Peñasco serpentinites plot within the superposition field of mid-ocean ridge (MOR) and suprasubduction zone (SSZ) peridotites. D)  $\text{Cr}^{\#}_{\text{r}}$  versus  $\text{Fe}^{2+}/(\text{Mg}^{2+} + \text{Fe}^{2+})$  diagram. Analyzed Cr-spinels plot within the Ocean Floor Peridotites and follow the Ca-Al trend described by Barnes and Roeder (2001).



**FIGURE 7.** A) Chondrite-normalized (Sun and McDonough, 1989) REE diagram for the Peñasco serpentinites. B) Primitive mantle normalized (Sun and McDonough, 1989) trace element diagram.

fluid involved in the metasomatism of serpentinites was a moderate to high temperature (350°–400°C), CO<sub>2</sub>-rich, mildly basic hydrothermal fluid related to the addition of Ca and Sr as has been interpreted by Akbulut *et al.* (2006) for serpentinites from the Mihaliçcik region (Turkey). Seawater could be a possible source for CO<sub>2</sub>, Ca and Sr. Nevertheless, the presence of listvenites bodies at serpentinite margins with quartz, serpentines, calcite as typical mineral assemblage suggests a multistage fluid influx, where a second type of fluid, which lead at least to quartz precipitation at serpentinite margins, also intervened. Akbulut *et al.* (2006) and Tsikouras *et al.* (2006) describe that the influx of SiO<sub>2</sub> can be driven by low-temperature (<250°C), highly-oxidized (high fO<sub>2</sub>), highly-acid (low pH), saline-rich fluids. The high fO<sub>2</sub> and fS<sub>2</sub> at decreasing temperatures also promotes the precipitation of magnetite and pyrite. Taking into account other examples throughout the world, the high rate of element mobilization during metasomatism should highlight the Precordillera listvenites as potential target for metallic mineralization.

### Comparative stratigraphy

The most accepted tectonic evolution of the Cuyania terrane initiates with rifting from Laurentia (Ouachita embayment) in the Neoproterozoic–early Cambrian, and continues with the development of a passive margin in latest early Cambrian to Early Ordovician, and collision with the Gondwana margin in the Middle–Late Ordovician (Astini *et al.*, 1995; Naipauer *et al.*, 2010; Thomas *et al.*, 2012; among others).

On the contrary, very little is known about the provenance and tectonic evolution of the Chilenia terrane. Metamorphic rocks (Las Yaretas Gneisses) exposed west of the Guarguaraz area yield U–Pb ages in zircons between 1060 and 1080 Ma (Ramos and Basei, 1997). Due to their age compatibility with Grenvillian rocks from Precordillera (Kay *et al.*, 1996; Rapela *et al.*, 2010), Ramos and Basei (1997) outlined two possible interpretations: i) Las Yaretas Gneisses belong to the Chilenia basement, and therefore,

Cuyania and Chilenia possibly derived from the same laurentian source or, ii) they represent imbricated slices of the basement of the Cuyania terrane emplaced into metasedimentary sequences located at Frontal Cordillera. Considering the last hypothesis as most likely, López de Azarevich *et al.* (2009) interpreted the Guarguaraz Complex as continental margin deposits of Vendian–early Cambrian age (~650–520Ma) developed along the western margin of the Cuyania terrane. Nevertheless, the two aforementioned hypotheses are still viable since no new compelling data are available. Therefore, the assignment of the Guarguaraz Complex to either the Cuyania or Chilenia terranes is currently a matter of debate. There is a general consensus that western Precordillera localities (Bonilla, Cortaderas, Peñasco) belong to the western margin of the Cuyania terrane (Ramos *et al.*, 1984, 1986, among others). The Bonilla Complex is assigned to a continental margin setting and has a maximum sedimentation age of *ca.* 592Ma (Gregori *et al.*, 2013), thus it can be considered temporarily equivalent to the Guarguaraz Complex, which yield a maximum sedimentation age of 555±8Ma (Willner *et al.*, 2008). To the north, in the Cortaderas and Peñasco areas, the Cortadera Complex is similar to the Guarguaraz and Bonilla complexes regarding their lithology, structure, metamorphic grade, and their mafic-ultramafic association. The Cortadera Complex sedimentation age is not yet constrained, but Davis *et al.* (2000) dated in 576±17Ma (U–Pb on zircon) a metabasaltic dike that intrudes phyllites, hence its time span could be extended to the late Neoproterozoic and be a temporal equivalent of the Bonilla and Guarguaraz complexes.

The Neoproterozoic basaltic magmatism that intrude the metasedimentary successions show a N- and E-MORB chemical and mantle isotopic signatures, with  $\epsilon_{\text{Nd}}$  between +3.8 and +8.2 (Davis *et al.*, 2000; López de Azarevich *et al.*, 2009). Basei *et al.* (1997) and López de Azarevich *et al.* (2009) estimated Sm–Nd model ages (TDM) between 0.57 and 1.62Ga for mafic bodies from the Guarguaraz area, which are interpreted to represent a Grenvillian mantle source for this basaltic magmatism.

Serpentinites from different complexes exhibit similar geochemical fingerprints. The Peñasco serpentinites show subhorizontal HREE slopes, which evidence that they were originated out of the garnet stability field, a feature also observed in the Guarguaraz serpentinites (López and Gregori, 2004; López de Azarevich *et al.*, 2009). Besides, Cr-spinel chemistry reveals that the Peñasco serpentinites show an intermediate signature between a mid-ocean ridge and an arc-related ophiolite. The presence of serpentinitized ultramafic rocks along the suture zone between the Cuyania and Chilenia terranes (Guarguaraz, Cortadera and Bonilla complexes) spatially associated to a N- and E-MORB basaltic magmatism and continental margin

metasedimentary successions evidence the development of an oceanic basin between the Cuyania and Chilenia terranes during the Neoproterozoic–early Paleozoic within an extensional regime associated to the Rodinia break-up.

Deformation associated with serpentinite obduction and tectonic emplacement related to the closure of the aforementioned oceanic basin generated faults, fractures and joints which constituted excellent pathways for hydrothermal fluids pervasive circulation. Serpentinization and listvenitization processes might have occurred during the opening and closure of the oceanic basin that separated the Chilenia and Cuyania terranes in Neoproterozoic–early Paleozoic times (Ramos *et al.*, 1984, 1986; Davis *et al.*, 1999, 2000; Willner *et al.*, 2011).

## CONCLUSIONS

In this contribution, we present the first geochemical whole-rock data set and the first chemical data set of Cr-spinels from the Cordón del Peñasco serpentinites.

Cr-spinels analyzed (picotite, pleonaste and spinel *sensu stricto*) show holly leaf shapes. They are chemically homogeneous and show mid-ocean ridge to arc affinities. They could have evolved from a mid-ocean ridge to an arc-related ophiolite related to the tectonic evolution of the Cuyania and Chilenia terranes.

Pseudomorphic textures formed by chrysotile/lizardite + magnetite ± brucite ± chlorite allow us to interpret that serpentinites derived from harzburgitic cumulates subjected to an ocean-floor hydrothermal alteration at temperatures  $\leq 300^\circ\text{C}$ .

Multistage metasomatic processes involving carbonatization and listvenitization also affected the serpentinites. Ca, Sr, LREE (specially Eu) and MREE enrichments were generated by a hydrothermal/serpentinite interaction at T below  $400^\circ\text{C}$ . The fluid was  $\text{CO}_2$  and complexing ion ( $\text{F}^-$ ,  $\text{Cl}^-$ ,  $\text{CO}_3^{2-}$ ,  $\text{PO}_4^{3-}$ )-rich, with mildly basic conditions for REE complexation.

The presence of quartz in listvenite mineralogy suggests the occurrence of a second type of fluid with low pH,  $\text{fO}_2$  and  $T \sim 250^\circ\text{C}$ .

An extensional regime associated to the Rodinia break-up may have led to the development of an oceanic basin between the Chilenia and Cuyania terranes at Neoproterozoic–early Paleozoic times evidenced by the presence of the metasomatized serpentinite bodies studied here and the occurrence of N- to E-MORB magmatism coeval with continental margin sedimentation of that age.

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