



Revista Mexicana de Física

ISSN: 0035-001X

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Sociedad Mexicana de Física A.C.

México

Félix, O.; Mondragón, A.; Mondragón, M.; Peinado, E.
Neutrino masses and mixings in a minimal S_3 -invariant extension of the standard model
Revista Mexicana de Física, vol. 52, núm. 4, noviembre, 2006, pp. 67-73
Sociedad Mexicana de Física A.C.
Distrito Federal, México

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Neutrino masses and mixings in a minimal S_3 -invariant extension of the standard model

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Recibido el 1 de febrero de 2006; aceptado el 7 de marzo de 2006

The mass matrices of the charged leptons and neutrinos, that had been derived in the framework of a Minimal S_3 -invariant Extension of the Standard Model, are reparametrized in terms of the masses of the charged leptons and neutrinos. Then, the neutrino mixing matrix V_{PMNS} is computed and we obtain explicit, analytic and exact expressions for the neutrino mixing angles as function of the masses of the charged leptons and neutrinos. By comparison with the experimental data on neutrino mixings, we obtain numerical values for the neutrino masses in good agreement with the experimental bounds extracted from neutrinoless double beta decay and precision cosmological observations.

Keywords: Flavor symmetries; quark and lepton masses and mixings; neutrino masses and mixings.

Las matrices de masas de los leptones cargados y los neutrinos, que previamente habíamos obtenido en una extensión Mínima S_3 -invariante del Modelo Standard, se reparametrizan en términos de las masas de los leptones cargados y los neutrinos. A partir de aquí, se calcula la matriz de mezclas de los neutrinos, V_{PMNS} , y se obtienen expresiones analíticas, explícitas y exactas para los ángulos de mezcla y las fases de Majorana como función de las masas de los leptones cargados y los neutrinos. Por comparación con los datos experimentales sobre mezclas de neutrinos, obtenemos valores numéricos de las masas de los neutrinos en buen acuerdo con las cotas experimentales extraídas de las medidas de la desintegración beta, la desintegración doble beta sin neutrinos y las observaciones cosmológicas de precisión.

Descriptores: Simetrías del sabor; masas y mezclas de quarks y leptones; masas y mezclas de neutrinos.

PACS: 11.30.Hv; 12.15.Ff; 14.60.Pq

1. Introduction

The discovery of neutrino masses and mixings marked a turning point in our understanding of nature and brought neutrino physics to the focus of attention of the particle, nuclear and astrophysics communities [1]. Recent neutrino oscillation observations and experiments have allowed the determination of the differences of the neutrino masses squared and the flavour mixing angles in the leptonic sector. The solar [2-5], atmospheric [6,7] and reactor [8,9] experiments produced the following results:

$$7.1 \times 10^{-5} (eV)^2 \leq \Delta^2 m_{12} \leq 8.9 \times 10^{-5} (eV)^2, \quad (1)$$

$$0.29 \leq \sin^2 \theta_{12} \leq 0.40, \quad (2)$$

$$1.4 \times 10^{-3} (eV)^2 \leq \Delta^2 m_{13} \leq 3.3 \times 10^{-3} (eV)^2, \quad (3)$$

$$0.34 \leq \sin^2 \theta_{23} \leq 0.68, \quad (4)$$

at 90% confidence level [10, 11]. The CHOOZ experiment [12] determined an upper bound for the flavour mixing angle between the first and the third generations:

$$\sin^2 \theta_{13} \leq 0.046. \quad (5)$$

Neutrino oscillation data are insensitive to the absolute value of neutrino masses and also to the fundamental issue of whether neutrinos are Dirac or Majorana particles. Hence, the importance of the upper bounds on neutrino masses provided by the searches that probe the neutrino mass values at rest: beta decay experiments [13], neutrinoless double beta decay [14] and precision cosmology [15].

On the theoretical side, the discovery of neutrino masses and mixings has also brought about important changes. In the Standard Model, the Higgs and Yukawa sectors, which are responsible for the generation of the masses of quarks and charged leptons, do not give mass to the neutrinos. Furthermore, the Yukawa sector of the Standard Model already has too many parameters whose values can not be determined from experiment. These two facts, taken together, point to the necessity and convenience of eliminating parameters and systematizing the observed hierarchies of masses and mixings, as well as the presence or absence of CP violating phases by means of a flavour or family symmetry under which the families transform in a non-trivial fashion. Such a flavour symmetry might be a continuous group or, more economically, a finite group.

In a recent paper, we argued that such a flavour symmetry, unbroken at the Fermi scale, is the permutational symmetry of three objects, S_3 , and introduced a Minimal S_3 -invariant Extension of the Standard Model [16]. In this model, we imposed S_3 as a fundamental symmetry in the matter sector. This assumption led us necessarily to extend the concept of flavour and generations to the Higgs sector. Hence, going to the irreducible representations of S_3 , we added to the Higgs $SU(2)_L$ doublet in the S_3 -singlet representation, two more Higgs $SU(2)_L$ doublets which can only belong to the two components of the S_3 -doublet representation. In this way, all the matter fields in the Minimal S_3 -invariant Extension of the Standard Model - Higgs, quark and lepton fields, including the right handed neutrino fields- belong to the three dimensional representation $\mathbf{1} \oplus \mathbf{2}$ of the permutational group

S_3 . The leptonic sector of the model was further constrained by an Abelian Z_2 symmetry.

The group S_3 [17-25] and the product groups $S_3 \times S_3$ [25-28] and $S_3 \times S_3 \times S_3$ [29,30] have been considered by many authors to explain successfully the hierarchical structure of quark masses and mixings in the Standard Model. However, in these works, the S_3 , $S_3 \times S_3$ and $S_3 \times S_3 \times S_3$ symmetries are explicitly broken at the Fermi scale to give mass to the lighter quarks and charged leptons, neutrinos are left massless. Some other interesting models based on the S_3 , S_4 and A_4 flavour symmetry groups, unbroken at the Fermi scale, have also been proposed [31-36], but in those models, equality of the number of fields and the irreducible representations is not obtained.

In this paper, we derive exact, explicit, analytic expression for the elements of the leptonic mixing matrix, V_{PMNS} , as functions of the masses of the charged leptons and the neutrinos. By comparison with the latest experimental data on neutrino mixings, we obtain numerical values for the neutrino masses in good agreement with the experimental bounds extracted from the precision cosmological observations and the neutrinoless double beta decay.

2. The Minimal S_3 -invariant Extension of the Standard Model

In the Standard Model analogous fermions in different generations have completely identical couplings to all gauge bosons of the strong, weak and electromagnetic interactions. Prior to the introduction of the Higgs boson and mass terms, the Lagrangian is chiral and invariant with respect to permutations of the left and right fermionic fields.

The six possible permutations of three objects (f_1, f_2, f_3) are elements of the permutational group S_3 . This is the discrete, non-Abelian group with the smallest number of elements. The three-dimensional real representation is not an irreducible representation of S_3 . It can be decomposed into the direct sum of a doublet f_D and a singlet f_s , where

$$f_s = \frac{1}{\sqrt{3}}(f_1 + f_2 + f_3),$$

$$f_D^T = \left(\frac{1}{\sqrt{2}}(f_1 - f_2), \frac{1}{\sqrt{6}}(f_1 + f_2 - 2f_3) \right). \quad (6)$$

The direct product of two doublets $\mathbf{p}_D^T = (p_{D1}, p_{D2})$ and $\mathbf{q}_D^T = (q_{D1}, q_{D2})$ may be decomposed into the direct sum of two singlets \mathbf{r}_s and $\mathbf{r}_{s'}$, and one doublet \mathbf{r}_D^T where

$$\mathbf{r}_s = p_{D1}q_{D1} + p_{D2}q_{D2}, \mathbf{r}_{s'} = p_{D1}q_{D2} - p_{D2}q_{D1}, \quad (7)$$

$$\mathbf{r}_D^T = (r_{D1}, r_{D2})$$

$$= (p_{D1}q_{D2} + p_{D2}q_{D1}, p_{D1}q_{D1} - p_{D2}q_{D2}). \quad (8)$$

The antisymmetric singlet $\mathbf{r}_{s'}$ is not invariant under S_3 .

Since the Standard Model has only one Higgs $SU(2)_L$ doublet, which can only be an S_3 singlet, it can only give

mass to the quark or charged lepton in the S_3 singlet representation, one in each family, without breaking the S_3 symmetry.

Hence, in order to impose S_3 as a fundamental symmetry, unbroken at the Fermi scale, we are led to extend the Higgs sector of the theory. The quark, lepton and Higgs fields are

$$Q^T = (u_L, d_L), u_R, d_R,$$

$$L^T = (\nu_L, e_L), e_R, \nu_R \text{ and } H, \quad (9)$$

in an obvious notation. All of these fields have three species, and we assume that each forms a reducible representation $\mathbf{1}_S \oplus \mathbf{2}$. The doublets carry capital indices I and J , which run from 1 to 2, and the singlets are denoted by $Q_3, u_{3R}, d_{3R}, L_3, e_{3R}, \nu_{3R}$ and H_S . Note that the subscript 3 denotes the singlet representation and not the third generation. The most general renormalizable Yukawa interactions are given by

$$\mathcal{L}_Y = \mathcal{L}_{Y_D} + \mathcal{L}_{Y_U} + \mathcal{L}_{Y_E} + \mathcal{L}_{Y_\nu}, \quad (10)$$

where

$$\mathcal{L}_{Y_D} = -Y_1^d \bar{Q}_I H_S d_{IR} - Y_3^d \bar{Q}_3 H_S d_{3R}$$

$$- Y_2^d [\bar{Q}_I \kappa_{IJ} H_1 d_{JR} + \bar{Q}_I \eta_{IJ} H_2 d_{JR}]$$

$$- Y_4^d \bar{Q}_3 H_I d_{IR} - Y_5^d \bar{Q}_I H_I d_{3R} + \text{h.c.}, \quad (11)$$

$$\mathcal{L}_{Y_U} = -Y_1^u \bar{Q}_I (i\sigma_2) H_S^* u_{IR} - Y_3^u \bar{Q}_3 (i\sigma_2) H_S^* u_{3R}$$

$$- Y_2^u [\bar{Q}_I \kappa_{IJ} (i\sigma_2) H_1^* u_{JR} + \eta \bar{Q}_I \eta_{IJ} (i\sigma_2) H_2^* u_{JR}]$$

$$- Y_4^u \bar{Q}_3 (i\sigma_2) H_I^* u_{IR} - Y_5^u \bar{Q}_I (i\sigma_2) H_I^* u_{3R} + \text{h.c.}, \quad (12)$$

$$\mathcal{L}_{Y_E} = -Y_1^e \bar{L}_I H_S e_{IR} - Y_3^e \bar{L}_3 H_S e_{3R}$$

$$- Y_2^e [\bar{L}_I \kappa_{IJ} H_1 e_{JR} + \bar{L}_I \eta_{IJ} H_2 e_{JR}]$$

$$- Y_4^e \bar{L}_3 H_I e_{IR} - Y_5^e \bar{L}_I H_I e_{3R} + \text{h.c.}, \quad (13)$$

$$\mathcal{L}_{Y_\nu} = -Y_1^\nu \bar{L}_I (i\sigma_2) H_S^* \nu_{IR} - Y_3^\nu \bar{L}_3 (i\sigma_2) H_S^* \nu_{3R}$$

$$- Y_2^\nu [\bar{L}_I \kappa_{IJ} (i\sigma_2) H_1^* \nu_{JR} + \bar{L}_I \eta_{IJ} (i\sigma_2) H_2^* \nu_{JR}]$$

$$- Y_4^\nu \bar{L}_3 (i\sigma_2) H_I^* \nu_{IR} - Y_5^\nu \bar{L}_I (i\sigma_2) H_I^* \nu_{3R} + \text{h.c.}, \quad (14)$$

and

$$\kappa = \begin{pmatrix} 0 & 1 \\ 1 & 0 \end{pmatrix} \text{ and } \eta = \begin{pmatrix} 1 & 0 \\ 0 & -1 \end{pmatrix}. \quad (15)$$

Furthermore, we add to the Lagrangian the Majorana mass terms for the right-handed neutrinos

$$\mathcal{L}_M = -M_1 \nu_{IR}^T C \nu_{IR} - M_3 \nu_{3R}^T C \nu_{3R}. \quad (16)$$

Due to the presence of three Higgs fields, the Higgs potential $V_H(H_S, H_D)$ is more complicated than that of the Standard Model. This potential was analyzed by Pakvasa and Sugawara [18] who found that in addition to the S_3 symmetry, it has a permutational symmetry S_2 : $H_1 \leftrightarrow H_2$, which is not a subgroup of the flavour group S_3 and an Abelian discrete symmetry that we will use for selection rules of the

Yukawa couplings in the leptonic sector. In this communication, we will assume that the vacuum respects the accidental S_2 symmetry of the Higgs potential and

$$\langle H_1 \rangle = \langle H_2 \rangle. \quad (17)$$

With these assumptions, the Yukawa interactions, Eqs. (11)-(14) yield mass matrices, for all fermions in the theory, of the general form

$$\mathbf{M} = \begin{pmatrix} \mu_1 + \mu_2 & \mu_2 & \mu_5 \\ \mu_2 & \mu_1 - \mu_2 & \mu_5 \\ \mu_4 & \mu_4 & \mu_3 \end{pmatrix}. \quad (18)$$

The Majorana mass for the left neutrinos ν_L will be obtained from the see-saw mechanism. The corresponding mass matrix is given by

$$\mathbf{M}_\nu = \mathbf{M}_{\nu_D} \tilde{\mathbf{M}}^{-1} (\mathbf{M}_{\nu_D})^T \quad (19)$$

where $\tilde{\mathbf{M}} = \text{diag}(M_1, M_1, M_3)$.

In principle, all entries in the mass matrices can be complex since there is no restriction coming from the flavour symmetry S_3 .

The mass matrices are diagonalized by bi-unitary transformations as

$$U_{d(u,e)L}^\dagger \mathbf{M}_{d(u,e)} U_{d(u,e)R} = \text{diag}(m_{d(u,e)}, m_{s(c,\mu)}, m_{b(t,\tau)}), \\ U_\nu^T \mathbf{M}_\nu U_\nu = \text{diag}(m_{\nu_1}, m_{\nu_2}, m_{\nu_3}). \quad (20)$$

The entries in the diagonal matrices may be complex, so the physical masses are their absolute values.

The mixing matrices are, by definition,

$$V_{CKM} = U_{uL}^\dagger U_{dL}, \quad V_{PMNS} = U_{eL}^\dagger U_\nu K, \quad (21)$$

where K is the diagonal matrix of the Majorana phases.

3. The mass matrices in the leptonic sector and Z_2 symmetry

A further reduction of the number of parameters in the leptonic sector may be achieved by means of an Abelian Z_2 symmetry. A possible set of charge assignments of Z_2 , compatible with the experimental data on masses and mixings in the leptonic sector is given in Table I

These Z_2 assignments forbid certain Yukawa couplings,

$$Y_1^e = Y_3^e = Y_1^\nu = Y_5^\nu = 0. \quad (22)$$

Therefore, the corresponding entries in the mass matrices vanish, i.e., $\mu_1^e = \mu_3^e = 0$ and $\mu_1^\nu = \mu_5^\nu = 0$.

TABLE I. Z_2 assignment in the leptonic sector.

-		+	
H_S, ν_{3R}		$H_I, L_3, L_I, e_{3R}, e_{IR}, \nu_{IR}$	

3.1. The mass matrix of the charged leptons

The mass matrix of the charged leptons takes the form

$$M_e = m_\tau \begin{pmatrix} \tilde{\mu}_2 & \tilde{\mu}_2 & \tilde{\mu}_5 \\ \tilde{\mu}_2 & -\tilde{\mu}_2 & \tilde{\mu}_5 \\ \tilde{\mu}_4 & \tilde{\mu}_4 & 0 \end{pmatrix}. \quad (23)$$

The unitary matrix U_{eL} that enters in the definition of the mixing matrix, V_{PMNS} , is calculated from

$$U_{eL}^\dagger M_e M_e^\dagger U_{eL} = \text{diag}(m_e^2, m_\mu^2, m_\tau^2), \quad (24)$$

where m_e, m_μ and m_τ are the masses of the charged leptons, and

$$M_e M_e^\dagger = m_\tau^2 \times \begin{pmatrix} 2|\tilde{\mu}_2|^2 + |\tilde{\mu}_5|^2 & |\tilde{\mu}_5|^2 & 2|\tilde{\mu}_2||\tilde{\mu}_4|e^{-i\delta_e} \\ |\tilde{\mu}_5|^2 & 2|\tilde{\mu}_2|^2 + |\tilde{\mu}_5|^2 & 0 \\ 2|\tilde{\mu}_2||\tilde{\mu}_4|e^{i\delta_e} & 0 & 2|\tilde{\mu}_4|^2 \end{pmatrix}, \quad (25)$$

Notice that this matrix has only one non-ignorable phase factor.

The parameters $|\tilde{\mu}_2|$, $|\tilde{\mu}_4|$ and $|\tilde{\mu}_5|$ may readily be expressed in terms of the charged lepton masses. From the invariants of $M_e M_e^\dagger$, we get the set of equations

$$\text{Tr}(M_e M_e^\dagger) = m_e^2 + m_\mu^2 + m_\tau^2 \\ = m_\tau^2 [4|\tilde{\mu}_2|^2 + 2(|\tilde{\mu}_4|^2 + |\tilde{\mu}_5|^2)], \quad (26)$$

$$\chi(M_e M_e^\dagger) = m_\tau^2(m_e^2 + m_\mu^2) + m_e^2 m_\mu^2 \\ = 4m_\tau^4 [|\tilde{\mu}_2|^4 + |\tilde{\mu}_2|^2(|\tilde{\mu}_4|^2 + |\tilde{\mu}_5|^2) + |\tilde{\mu}_4|^2|\tilde{\mu}_5|^2], \quad (27)$$

$$\det(M_e M_e^\dagger) = m_e^2 m_\mu^2 m_\tau^2 = 4m_\tau^6 |\tilde{\mu}_2|^2 |\tilde{\mu}_4|^2 |\tilde{\mu}_5|^2, \quad (28)$$

where $\chi(M_e M_e^\dagger) = \frac{1}{2} [(Tr(M_e M_e^\dagger))^2 - Tr(M_e M_e^\dagger)^2]$.

Solving these equations for $|\tilde{\mu}_2|^2$, $|\tilde{\mu}_4|^2$ and $|\tilde{\mu}_5|^2$, we obtain

$$|\tilde{\mu}_2|^2 = \frac{1}{2} \frac{m_e^2 + m_\mu^2}{m_\tau^2} - \frac{m_e^2 m_\mu^2}{m_\tau^2(m_e^2 + m_\mu^2)} + \beta, \quad (29)$$

in this expression, β is the smallest solution of the equation

$$\beta^3 - \frac{1}{2} \left(1 - 2y + 6\frac{z}{y} \right) \beta^2 \\ - \frac{1}{4} \left(y - y^2 - 4\frac{z}{y} + 7z - 12\frac{z^2}{y^2} \right) \beta \\ - \frac{1}{8} yz - \frac{1}{2} \frac{z^2}{y^2} + \frac{3}{4} \frac{z^2}{y} - \frac{z^3}{y^3} = 0. \quad (30)$$

where $y = (m_e^2 + m_\mu^2)/m_\tau^2$ and $z = m_\mu^2 m_e^2 / m_\tau^4$.

A good, order of magnitude, estimation for β is obtained from (30)

$$\beta \approx -\frac{m_\mu^2 m_e^2}{2m_\tau^2(m_\tau^2 - (m_\mu^2 + m_e^2))}. \quad (31)$$

The parameters $|\tilde{\mu}_4|^2$ and $|\tilde{\mu}_5|^2$ are, then, readily expressed in terms of $|\tilde{\mu}_2|^2$,

$$|\tilde{\mu}_{4,5}|^2 = \frac{1}{4} \left(1 - \frac{m_\mu^2 + m_e^2}{m_\tau^2} + 4 \frac{m_e^2 m_\mu^2}{m_\tau^2 (m_e^2 + m_\mu^2)} - 4\beta \right) \pm \sqrt{\left(1 - \frac{m_\mu^2 + m_e^2}{m_\tau^2} + 4 \frac{m_e^2 m_\mu^2}{m_\tau^2 (m_e^2 + m_\mu^2)} - 4\beta \right)^2 - \frac{m_\mu^2 m_e^2}{m_\tau^4} \frac{1}{|\tilde{\mu}_2|^2}}. \quad (32)$$

Once $M_e M_e^\dagger$ has been reparametrized in terms of the charged lepton masses, it is straightforward to compute U_{eL} also as a function of the lepton masses. Here, in order to avoid a clumsy notation, we will write the result to order of $(m_\mu m_e / m_\tau^2)^2$.

$$M_e \approx m_\tau \frac{1}{\sqrt{2}} \begin{pmatrix} \frac{\sqrt{m_\mu^2 + m_e^2}}{m_\tau} \sqrt{1 - 2 \frac{m_\mu^2 m_e^2}{(m_\mu^2 + m_e^2) m_\tau^2}} & \frac{\sqrt{m_\mu^2 + m_e^2}}{m_\tau} \sqrt{1 - 2 \frac{m_\mu^2 m_e^2}{(m_\mu^2 + m_e^2) m_\tau^2}} & \sqrt{1 - \frac{m_\mu^2 + m_e^2}{m_\tau^2} + 4 \frac{m_e^2 m_\mu^2}{m_\tau^2 (m_\mu^2 + m_e^2)}} \\ \frac{\sqrt{m_\mu^2 + m_e^2}}{m_\tau} \sqrt{1 - 2 \frac{m_\mu^2 m_e^2}{(m_\mu^2 + m_e^2) m_\tau^2}} & -\frac{\sqrt{m_\mu^2 + m_e^2}}{m_\tau} \sqrt{1 - 2 \frac{m_\mu^2 m_e^2}{(m_\mu^2 + m_e^2) m_\tau^2}} & \sqrt{1 - \frac{m_\mu^2 + m_e^2}{m_\tau^2} + 4 \frac{m_e^2 m_\mu^2}{m_\tau^2 (m_\mu^2 + m_e^2)}} \\ \frac{m_\mu m_e}{m_\tau \sqrt{m_\mu^2 - m_e^2}} e^{i\delta_e} & \frac{m_\mu m_e}{m_\tau \sqrt{m_\mu^2 - m_e^2}} e^{i\delta_e} & 0 \end{pmatrix} \quad (33)$$

and

$$U_{eL} \approx \begin{pmatrix} \frac{1}{\sqrt{2}} \frac{\frac{m_e}{m_\mu}}{\sqrt{1 - \left(\frac{m_e}{m_\mu}\right)^2}} & \frac{1}{\sqrt{2}} \frac{1}{\sqrt{1 + \left(\frac{m_e}{m_\mu}\right)^2}} & \frac{1}{\sqrt{2}} \frac{1}{\sqrt{1 + \frac{m_e m_\mu}{m_\tau^2}}} \\ -\frac{1}{\sqrt{2}} \frac{\frac{m_e}{m_\mu}}{\sqrt{1 - \left(\frac{m_e}{m_\mu}\right)^2}} & -\frac{1}{\sqrt{2}} \frac{1}{\sqrt{1 + \left(\frac{m_e}{m_\mu}\right)^2}} & \frac{1}{\sqrt{2}} \frac{1}{\sqrt{1 + \frac{m_e m_\mu}{m_\tau^2}}} \\ \frac{\sqrt{1 - 2 \left(\frac{m_e}{m_\mu}\right)^2}}{\sqrt{1 - \left(\frac{m_e}{m_\mu}\right)^2}} e^{i\delta_e} & \frac{\frac{m_e}{m_\mu} e^{i\delta_e}}{\sqrt{1 + \left(\frac{m_e}{m_\mu}\right)^2}} & \frac{\frac{m_e m_\mu}{m_\tau^2} e^{i\delta_e}}{\sqrt{1 + \frac{m_e m_\mu}{m_\tau^2}}} \end{pmatrix} \quad (34)$$

4. The mass matrix of the neutrinos

According with the Z_2 selection rule eq. (22), the mass matrix of the Dirac neutrino takes the form

$$\mathbf{M} = \begin{pmatrix} \mu_2^\nu & \mu_2^\nu & 0 \\ \mu_2^\nu & -\mu_2^\nu & 0 \\ \mu_4^\nu & \mu_4^\nu & \mu_3^\nu \end{pmatrix}. \quad (35)$$

Then, the mass matrix for the left-handed Majorana neutrinos obtained from the see-saw mechanism as

$$M_\nu = M_{\nu_D} \tilde{\mathbf{M}}^{-1} (M_{\nu_D})^T \quad (36)$$

$$= \begin{pmatrix} 2(\rho_2^\nu)^2 & 0 & 2\rho_2^\nu \rho_4^\nu \\ 0 & 2(\rho_2^\nu)^2 & 0 \\ 2\rho_2^\nu \rho_4^\nu & 0 & 2(\rho_4^\nu)^2 + (\rho_3^\nu)^2 \end{pmatrix}, \quad (37)$$

where $\rho_2^\nu = (\mu_2^\nu) / M_1^{1/2}$, $\rho_4^\nu = (\mu_4^\nu) / M_1^{1/2}$ and $\rho_3^\nu = (\mu_3^\nu) / M_3^{1/2}$; M_1 and M_3 are the masses of the right handed neutrinos appearing in (16).

The non-Hermitian, complex, symmetric matrix M_ν may be brought to a diagonal form by a bi-unitary transformation, as

$$U_\nu^T M_\nu U_\nu = \text{diag}(|m_{\nu_1}| e^{i\phi_1}, |m_{\nu_2}| e^{i\phi_2}, |m_{\nu_3}| e^{i\phi_\nu}). \quad (38)$$

In order to compute U_ν , we notice that M_ν may be brought to a block diagonal form by a permutation of the second and third rows and columns

$$S_{23} M_\nu S_{23} = \begin{pmatrix} 2(\rho_2^\nu)^2 & 2\rho_2^\nu \rho_4^\nu & 0 \\ 2\rho_2^\nu \rho_4^\nu & 2(\rho_4^\nu)^2 + (\rho_3^\nu)^2 & 0 \\ 0 & 0 & 2(\rho_2^\nu)^2 \end{pmatrix}, \quad (39)$$

where

$$S_{23} = \begin{pmatrix} 1 & 0 & 0 \\ 0 & 0 & 1 \\ 0 & 1 & 0 \end{pmatrix}.$$

The entry in the lower right corner of the matrix in the right hand side of eq. (39), is already diagonal and may be identified with $m_{\nu_3} e^{i\phi_\nu}$; $|m_{\nu_3}|$ is the physical mass.

Then, the 2×2 block in the upper left hand corner may be diagonalized by a 2×2 bi-unitary transformation as

$$U_{2 \times 2}^T M_{2 \times 2} U_{2 \times 2} = \text{diag}(|m_{\nu_1}| e^{i\phi_1}, |m_{\nu_2}| e^{i\phi_2}).$$

$U_{2 \times 2}$ is computed from

$$U_{2 \times 2}^\dagger M_{2 \times 2} M_{2 \times 2}^\dagger U_{2 \times 2} = \text{diag}(|m_{\nu_1}|^2, |m_{\nu_2}|^2).$$

Since the 2×2 Hermitian matrix $M_{2 \times 2} M_{2 \times 2}^\dagger$ has only one phase factor, the unitary $U_{2 \times 2}$ matrix may be written as

$$U_{2 \times 2} = \begin{pmatrix} \sin \eta & \cos \eta \\ -\cos \eta e^{i\delta_\nu} & \sin \eta e^{i\delta_\nu} \end{pmatrix}. \quad (40)$$

Explicit expressions for the entries of $M_{2 \times 2}$ and $\sin \eta$, $\cos \eta$ as functions of the complex masses m_{ν_1} , m_{ν_2} and m_{ν_3} are obtained when the expression (40) for $U_{2 \times 2}$ is substituted in eq (38) and we require that this equation be satisfied as an identity. The result is

$$\cos \eta = \sqrt{\frac{m_{\nu_3} - m_{\nu_1}}{m_{\nu_2} - m_{\nu_1}}}, \quad \sin \eta = \sqrt{\frac{m_{\nu_2} - m_{\nu_3}}{m_{\nu_2} - m_{\nu_1}}} \quad (41)$$

and

$$M_\nu = \begin{pmatrix} m_{\nu_3} & 0 & \sqrt{(m_{\nu_3} - m_{\nu_1})(m_{\nu_2} - m_{\nu_3})} e^{-i\delta_\nu} \\ 0 & m_{\nu_3} & 0 \\ \sqrt{(m_{\nu_3} - m_{\nu_1})(m_{\nu_2} - m_{\nu_3})} e^{-i\delta_\nu} & 0 & (m_{\nu_1} + m_{\nu_2} - m_{\nu_3}) e^{-2i\delta_\nu} \end{pmatrix}. \quad (43)$$

$$U_\nu = \begin{pmatrix} \sqrt{\frac{m_{\nu_2} - m_{\nu_3}}{m_{\nu_2} - m_{\nu_1}}} & \sqrt{\frac{m_{\nu_3} - m_{\nu_1}}{m_{\nu_2} - m_{\nu_1}}} & 0 \\ 0 & 0 & 1 \\ -\sqrt{\frac{m_{\nu_3} - m_{\nu_1}}{m_{\nu_2} - m_{\nu_1}}} e^{i\delta_\nu} & \sqrt{\frac{m_{\nu_2} - m_{\nu_3}}{m_{\nu_2} - m_{\nu_1}}} e^{i\delta_\nu} & 0 \end{pmatrix}. \quad (44)$$

The only free parameters in these matrices, other than the neutrino masses, are the phase ϕ_ν , implicit in m_{ν_1} , m_{ν_2} and m_{ν_3} , and the Dirac phase δ_ν .

4.1. The neutrino mixing matrix

The neutrino mixing matrix V_{PMNS} , in the standard form advocated by the *PDG* [37], is obtained by taking the product $U_{eL}^\dagger U_\nu K$ and making an appropriate transformation of phases. Writing the resulting expression to the same approximation as in eq. (33), we get

$$V_{PMNS} \approx \begin{pmatrix} \frac{1}{\sqrt{2}} \frac{x}{\sqrt{1-x^2}} \sin \eta + \frac{\sqrt{1-2x^2}}{\sqrt{1-x^2}} \cos \eta & \frac{1}{\sqrt{2}} \frac{x}{\sqrt{1-x^2}} \cos \eta - \frac{\sqrt{1-2x^2}}{\sqrt{1-x^2}} \sin \eta & -\frac{1}{\sqrt{2}} \frac{x}{\sqrt{1-x^2}} e^{-i\delta} \\ \frac{1}{\sqrt{2}} \frac{1}{\sqrt{1+x^2}} \sin \eta - \frac{x}{\sqrt{1-x^2}} \cos \eta e^{i\delta} & \frac{1}{\sqrt{2}} \frac{1}{\sqrt{1+x^2}} \cos \eta + \frac{x}{\sqrt{1-x^2}} \sin \eta e^{i\delta} & -\frac{1}{\sqrt{2}} \frac{1+2\frac{z}{y(1-y)}}{\sqrt{1+x^2}} \\ \frac{1}{\sqrt{2}} \frac{1}{\sqrt{1+\sqrt{z}}} \sin \eta - \frac{\sqrt{z}}{\sqrt{1+\sqrt{z}}} \cos \eta e^{i\delta} & \frac{1}{\sqrt{2}} \frac{1}{\sqrt{1+\sqrt{z}}} \cos \eta + \frac{\sqrt{z}}{\sqrt{1+\sqrt{z}}} \sin \eta e^{i\delta} & \frac{1}{\sqrt{2}} \frac{1}{\sqrt{1+\sqrt{z}}} \end{pmatrix} K. \quad (45)$$

where $\cos \eta$ and $\sin \eta$ are given in eq. (41), $x = m_e/m_\mu$, $\delta = \delta_\nu - \delta_e$ and K is the diagonal matrix of the Majorana phases, $K = \text{diag}(1, e^{i\alpha}, e^{i\beta})$.

A comparison of this expression with the standard parametrization allowed us to derive expressions for the mixing angles in terms of the charged lepton and neutrino masses

$$\begin{aligned} \sin \theta_{13} &\approx \frac{\frac{1}{\sqrt{2}} \frac{m_e}{m_\mu}}{\sqrt{1 - \left(\frac{m_e}{m_\mu}\right)^2}}, \\ \sin \theta_{23} &\approx -\frac{1}{\sqrt{2}} \frac{\sqrt{1 - \left(\frac{m_e}{m_\mu}\right)^2}}{\sqrt{1 - \frac{1}{2} \left(\frac{m_e}{m_\mu}\right)^2}} \end{aligned} \quad (46)$$

The unitarity of $U_{2 \times 2}$ constrains $\sin \eta$ to be real and $|\sin \eta| \leq 1$. This condition fixes the phases ϕ_1 and ϕ_2 as

$$|m_{\nu_1}| \sin \phi_1 = |m_{\nu_2}| \sin \phi_2 = |m_{\nu_3}| \sin \phi_\nu. \quad (42)$$

The real phase δ_ν appearing in eq. (40) is not constrained by the unitarity of U_ν .

By doing a second permutation of the second and third rows and columns in eq. (39) and writing the entries in M_ν as functions of the complex masses m_{ν_1} , m_{ν_2} and m_{ν_3} , we find

and

$$\begin{aligned} \tan \theta_{12} &= -\sqrt{\frac{m_{\nu_2} - m_{\nu_3}}{m_{\nu_3} - m_{\nu_1}}} \\ &\times \left(\frac{\sqrt{1-2x^2} - \frac{1}{\sqrt{2}} x \sqrt{\frac{m_{\nu_3} - m_{\nu_1}}{m_{\nu_2} - m_{\nu_3}}}}{\sqrt{1-2x^2} + \frac{1}{\sqrt{2}} x \sqrt{\frac{m_{\nu_2} - m_{\nu_3}}{m_{\nu_3} - m_{\nu_1}}}} \right) \end{aligned} \quad (47)$$

The dependence of $\tan \theta_{12}$ on the phase ϕ_ν and the physical masses of the neutrinos is made explicit with the help of the unitarity constraint on U_ν , eq (42), we get

$$\begin{aligned} \tan^2 \eta &= \frac{m_{\nu_2} - m_{\nu_3}}{m_{\nu_3} - m_{\nu_1}} \\ &= -\frac{(|m_{\nu_2}|^2 - |m_{\nu_3}|^2 \sin^2 \phi_\nu)^{1/2} - |m_{\nu_3}| |\cos \phi_\nu|}{(|m_{\nu_1}|^2 - |m_{\nu_3}|^2 \sin^2 \phi_\nu)^{1/2} + |m_{\nu_3}| |\cos \phi_\nu|}. \end{aligned} \quad (48)$$

Similarly, the Majorana phases are given by

$$\begin{aligned}\sin 2\alpha = \sin(\phi_1 - \phi_2) &= \pm \frac{|m_{\nu_3}| \sin \phi_\nu}{|m_{\nu_1}| |m_{\nu_2}|} \times \left(\sqrt{|m_{\nu_2}|^2 - |m_{\nu_3}|^2 \sin^2 \phi_\nu} + \sqrt{|m_{\nu_1}|^2 - |m_{\nu_3}|^2 \sin^2 \phi_\nu} \right) \\ \sin 2\beta = \sin(\phi_1 - \phi_\nu) &= \pm \frac{\sin \phi_\nu}{|m_{\nu_1}|} \left(|m_{\nu_3}| \sqrt{1 - \sin^2 \phi_\nu} + \sqrt{|m_{\nu_1}|^2 - |m_{\nu_3}|^2 \sin^2 \phi_\nu} \right)\end{aligned}\quad (49)$$

A more complete and detailed discussion of the Majorana phases in the neutrino mixing matrix V_{PMNS} obtained in our model is given in J. Kubo [38].

In this model, $\sin^2 \theta_{12}$ and $\sin^2 \theta_{23}$ are determined by the masses of the charged leptons in very good agreement with the experimental values [10, 11, 39],

$$(\sin^2 \theta_{13})^{th} = 1.1 \times 10^{-5}, \quad (\sin^2 \theta_{13})^{exp} \leq 0.046,$$

and

$$(\sin^2 \theta_{23})^{th} = 0.49 \times 10^{-5}, \quad (\sin^2 \theta_{23})^{exp} = 0.5_{-0.05}^{+0.06}.$$

In the present model, the experimental restriction $|\Delta m_{12}^2| < |\Delta m_{13}^2|$ implies an inverted neutrino mass spectrum, $|m_{\nu_3}| < |m_{\nu_1}| < |m_{\nu_2}|$ [16].

As seen from Eqs. (47) and (48), the solar mixing angle is sensitive to the neutrino mass differences and the phase ϕ_ν but is only very weakly sensitive to the charged lepton masses. Writing the neutrino mass differences $m_{\nu_i} - m_{\nu_j}$ in terms of the differences of the mass squared and one of the neutrino masses, say $|m_{\nu_2}|$, from our previous expressions, (47) and (48), we obtain

$$\begin{aligned}\frac{m_{\nu_2}^2}{\Delta m_{13}^2} &= \frac{1 + 2t_{12}^2 + t_{12}^4 - rt_{12}}{4t_{12}^2(1 + t_{12}^2)(1 + t_{12}^2 - rt_{12}^2)\cos^2 \phi_\nu} \\ &\quad - \tan^2 \phi_\nu + O(x^2) \\ &\approx \frac{1}{\sin^2 2\theta_{12} \cos^2 \phi_\nu} - \tan^2 \phi_\nu \text{ for } r \ll 1,\end{aligned}\quad (50)$$

where $t_{12} = \tan \theta_{12}$ and $r = \Delta m_{21}^2 / \Delta m_{13}^2$.

The mass $|m_{\nu_2}|$ assumes its minimal value when $\sin \phi_\nu$ vanishes, then

$$|m_{\nu_2}| \approx \frac{\sqrt{\Delta m_{13}^2}}{\sin 2\theta_{12}}. \quad (51)$$

Hence, we find

$$\begin{aligned}|m_{\nu_2}| &\approx 0.0507 eV, \\ |m_{\nu_1}| &\approx 0.0499 eV \\ |m_{\nu_3}| &\approx 0.0193 eV\end{aligned}\quad (52)$$

where we used the values $\Delta m_{13}^2 = 2.2_{-0.27}^{+0.37} \times 10^{-3} eV^2$ and $\sin^2 \theta_{12} = 0.31_{-0.03}^{+0.02}$ taken from M. Maltoni *et al.* [10], T. Schwetz [11] and G.L. Fogli *et al.* [39].

With those values for the neutrino masses we compute the effective electron neutrino mass m_β ,

$$m_\beta = \left[\sum_i |U_{ei}|^2 m_{\nu_i}^2 \right]^{\frac{1}{2}} = 0.0502 eV, \quad (53)$$

well below the upper bound $m_\beta < 1.8 eV$ coming from the tritium β -decay experiments [13, 39, 40].

5. Conclusions

By introducing three Higgs fields that are $SU(2)_L$ doublets in the theory, we extended the concept of flavour and generations to the Higgs sector and formulated a Minimal S_3 -invariant Extension of the Standard Model [16]. A defined structure of the Yukawa couplings is obtained which permits the calculation of mass and mixing matrices for quarks and leptons in a unified way. A further reduction of redundant parameters is achieved in the leptonic sector by introducing a Z_2 symmetry. The flavour symmetry group $Z_2 \times S_3$ relates the mass spectrum and mixings. This allowed us to compute the neutrino mixing matrix explicitly in terms of the masses of charged leptons and neutrinos. In this model, the magnitudes of the three mixing angles are determined by the interplay of the flavour $S_3 \times Z_2$ symmetry, the see-saw mechanism and the lepton mass hierarchy. We also found that V_{PMNS} has one Dirac and two Majorana CP violating phases. The numerical values of the mixing angles are determined by the masses of the charged leptons only in very good agreement with the experiment. The solar mixing angle θ_{12} is almost insensitive to the values of the masses of the charged leptons, but its experimental value allowed us to fix the scale and origin of the neutrino mass spectrum which has an inverted hierarchy with the values $|m_{\nu_2}| = 0.0507 eV$, $|m_{\nu_1}| = 0.0499 eV$ and $|m_{\nu_3}| = 0.0193 eV$.

Acknowledgments

This work was partially supported by CONACYT México under contract No 40162-F and by DGAPA-UNAM under contract PAPIIT-IN116202-3.

1. C. K. Jung, C. Mc Grew, T. Kajita and T. Mann, *Annu. Rev. Nucl. Part. Sci.*, **51** (2001) 451.
2. M. Altmann *et al.* [GNO collaboration], *Phys. Lett. B* **616**, (2005), 174.
3. M. B. Smy *et al.* [SK collaboration], *Phys. Rev. D* **69** (2004), 011104.
4. Q.R. Ahmad *et al.* [SNO collaboration], *Phys. Rev. Lett.* **89** (2002) 011301.
5. B. Aharmim *et al.* [SNO collaboration], arxiv: nucl-ex/0502021.
6. S. Fukuda *et al.* [SK collaboration] *Phys. Lett. B* **539** (2002) 179.
7. Y. Ashie *et al.*, *Phys. Rev. Lett.* **93** (2004) 101801; [hep-ex/0501064].
8. C. Bemporad, G. Gratta and P. Vogel, *Rev. Mod. Phys.*, **74** (2002) 297.
9. T. Araki *et al.* (KamLAND collaboration), *Phys. Rev. Lett.* **94**, (2005) 081801.
10. M. Maltoni, T. Schwetz, M.A. Tórtola and J.W.F. Valle, *New J. Phys.* **6** (2004) 122.
11. T. Schwetz, “Neutrino oscillations: Current status and prospects”, arxiv: hep-ph/0510331.
12. M. Apollonio *et al.* [CHOOZ Collaboration], *Eur. Phys. J.* (2003) 331.
13. K. Eitel in “Neutrino 2004”, 21st International Conference on Neutrino Physics and Astrophysics (Paris, France 2004) Ed. J. Dumarchey, Th. Patyak and F. Vanucci. *Nucl. Phys. B (Proc Suppl.)* **143**, (2005) 3.
14. S. R. Eliot and J. Engel, *J. Phys. G* **30** (2004) R183.
15. O. Elgaroy, and O. Lahav, *New J. Phys.* **7** (2005) 61.
16. Kubo J, Mondragón A, Mondragón M, Rodríguez-Jáuregui E. *Prog. Theor. Phys.* **109**, (2003), 795.
17. H. Fritzsch *Phys. Lett.* **B70**, (1977), 436.
18. S. Pakvasa and H. Sugawara, *Phys. Lett.* **73B** (1978), 61.
19. H. Fritzsch, *Phys. Lett.* **B73**, (1978), 317.
20. H. Harari, H. Haut, J. Weyers, *Phys. Lett.* **B78** (1978), 459.
21. H. Fritzsch, *Nucl. Phys. B* **155**, (1979), 189.
22. Y. Yamanaka, S. Pakvasa and H. Sugawara, *Phys. Rev.* **D25** (1982), 1895.
23. P. Kaus and S. Meshkov, *Phys. Rev.* **D42** (1990), 1863.
24. H. Fritzsch and J.P. Plankl, *Phys. Lett.* **B237** (1990), 451.
25. P.F. Harrison and W.G. Scott, *Phys. Lett.* **B333** (1994), 471.
26. A. Mondragón and E. Rodríguez-Jáuregui, *Phys. Rev.* **D59** (1999), 093009.
27. A. Mondragón and E. Rodríguez-Jáuregui, *Phys. Rev.* **D61** (2000), 113002.
28. For a review see H. Fritzsch and Z.Z. Xing, *Prog. Part. Nucl. Phys.* **45** (2000) 1.
29. L.J. Hall and H. Murayama, *Phys. Rev. Lett.* **75** (1995), 3985.
30. C.D. Carone, L.J. Hall and H. Murayama, *Phys. Rev.* **D53** (1996), 6282.
31. Y. Koide, *Phys. Rev.* **D60** (1999), 077301.
32. E. Ma, *Phys. Rev.* **D44** (1991), 587.
33. E. Ma, *Mod. Phys. Lett. A* **17** (2002), 627; *ibid* **A17** (2002), 2361.
34. K.S. Babu, E. Ma and J.W.F. Valle, *Phys. Lett.* **B552** (2003), 207.
35. S.-L. Chen, M. Frigerio and E. Ma, *Phys. Rev.* **D70**, 073008 (2004); Erratum: *ibid* **D70** (2004), 079905.
36. W. Grimus and Lavoura J. *High Energy Phys.* **08** (2005) 013.
37. W-M Yao *et al* 2006 *J. Phys. G: Nucl. Part. Phys.* **33** 1.
38. J. Kubo, *Phys. Lett.* **B578**, (2004), 156; Erratum: *ibid* **619** (2005) 387.
39. G.L. Fogli, *et al.*, arxiv: hep-ph/0506083.
40. G. L. Fogli, E. Lisi, A. Marrone, A. Melchiorri, A. Palazzo, P. Serra and J. Silk, *Phys. Rev.* **D70** (2004), 113003.