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## UNVEILING THE STRUCTURE OF THE PLANETARY NEBULA M 2-48

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The PN M2-48 is formed by three main structures, namely, a bipolar central region (CR), a set of knots tracing a semicircular shell surrounding CR, and two symmetric bow-shocks. CR shows a kinematic structure corresponding to a bipolar shell, with an expansion velocity of  $\simeq 50\,\mathrm{km\,s^{-1}}$ . The semicircular shell appears to be expanding at  $\simeq 20\,\mathrm{km\,s^{-1}}$ , except in the regions aligned with the bow-shocks, which are interpreted as jet-shell interaction zones at  $\simeq 100\,\mathrm{km\,s^{-1}}$ . Finally, the bow-shocks have uncorrected velocities of  $\simeq 80\,\mathrm{km\,s^{-1}}$ . An inclination angle of  $10^{\circ}$  with respect to the plane of the sky is estimated using simple bow shock models.

High-dispersion (HD) spectroscopy was obtained in 1999, June 29 and 30, with the spectrometer MES (Meaburn et al. 1984) attached to the 2.1-m OAN telescope. A 90 Å bandwidth filter was used to isolate the 87th order, containing H $\alpha$  and [N II]  $\lambda\lambda6548$ , 6584 lines.

In Figure 1, the slit positions A-F are shown against a contour map of a [N II] 6584 image. The 150  $\mu \rm m$  wide (=  $10\,\rm km\,s^{-1})$  slit was oriented East-West (A-D) and North-South (E-F). The exposure times were of 1800 s for each slit position. The spectra were wavelength calibrated to an accuracy of  $\pm 1\,\rm km\,s^{-1}.$ 

An array of [N II] 6584 Å position-velocity (PV) maps, formed by the HD spectra (slits A-D), is shown in Figure 2. The complex kinematics of the main components of M 4-18 is evident, even revealing some features which were not detected previously.

The main results of our work are:

- The bright central bipolar region presents a heliocentric expansion velocity of 50 kms<sup>-1</sup>.
- 2. We propose that the knots around the bipolar core are forming two slow expanding arcs.
- 3. We confirm that the structure s1 is the interaction of an outflow with the east arc.

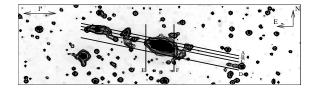


Fig. 1. The slit positions A-F are marked against a grayscale representation and contours of the [N II] emission of M 2-48.

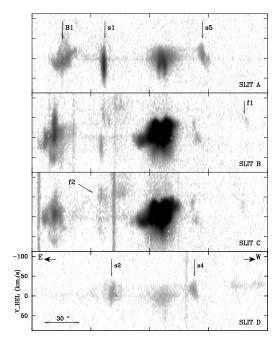


Fig. 2. Gray-scale representation of the PV arrays of [N II] 6584 Å along slits A-D. The different features are labeled following Vázquez et al. (2000).

4. We detect a bow-shock, B1, at  $\approx 2'$  of the bipolar core. The symmetrical counterpart is only marginally detected.

## REFERENCES

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